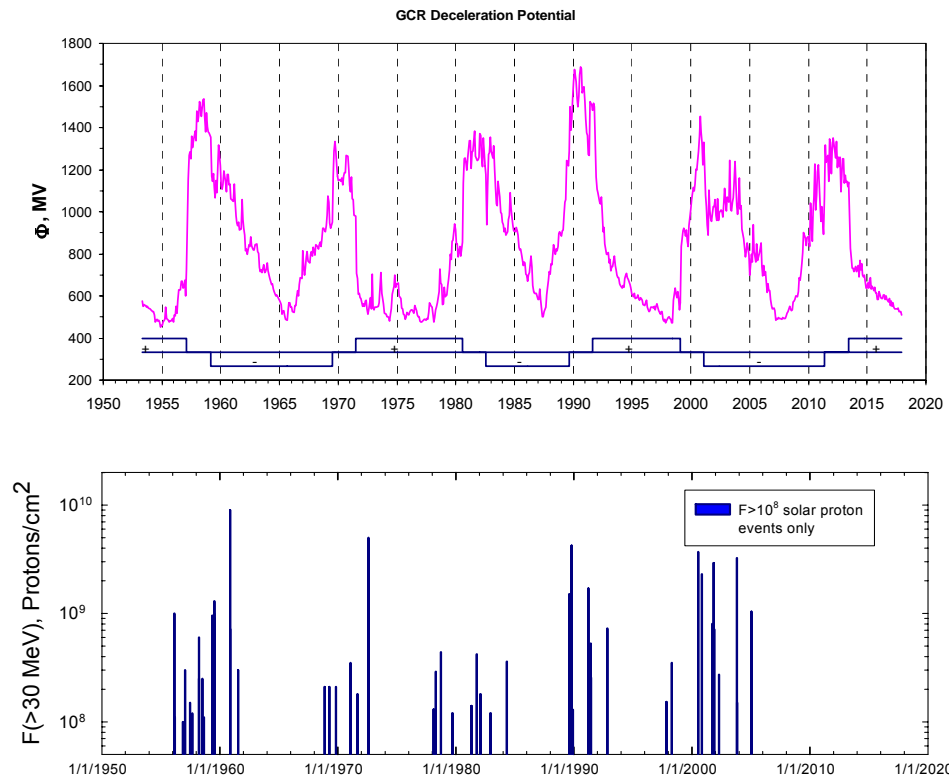


Example Solar Proton Event Data

NASA JSC
August 30, 2006

Times of Occurrence of Large SPE's



Modern Era (1956-2005)

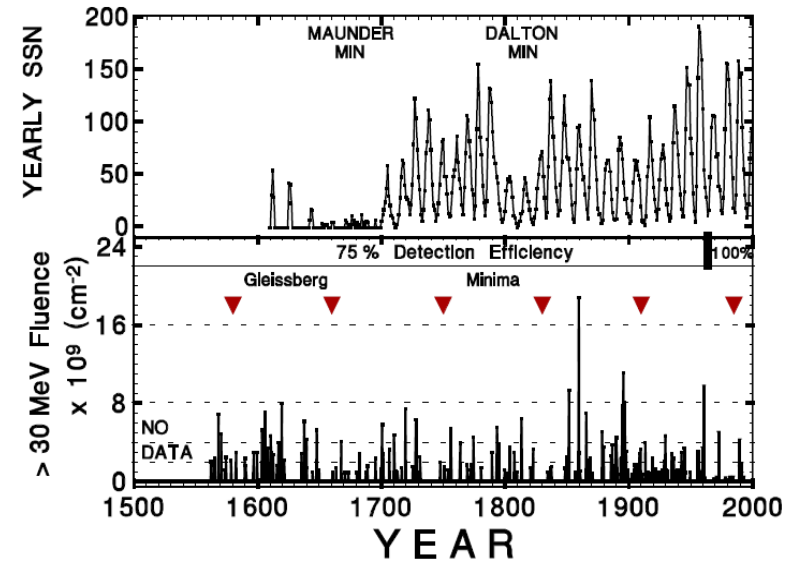
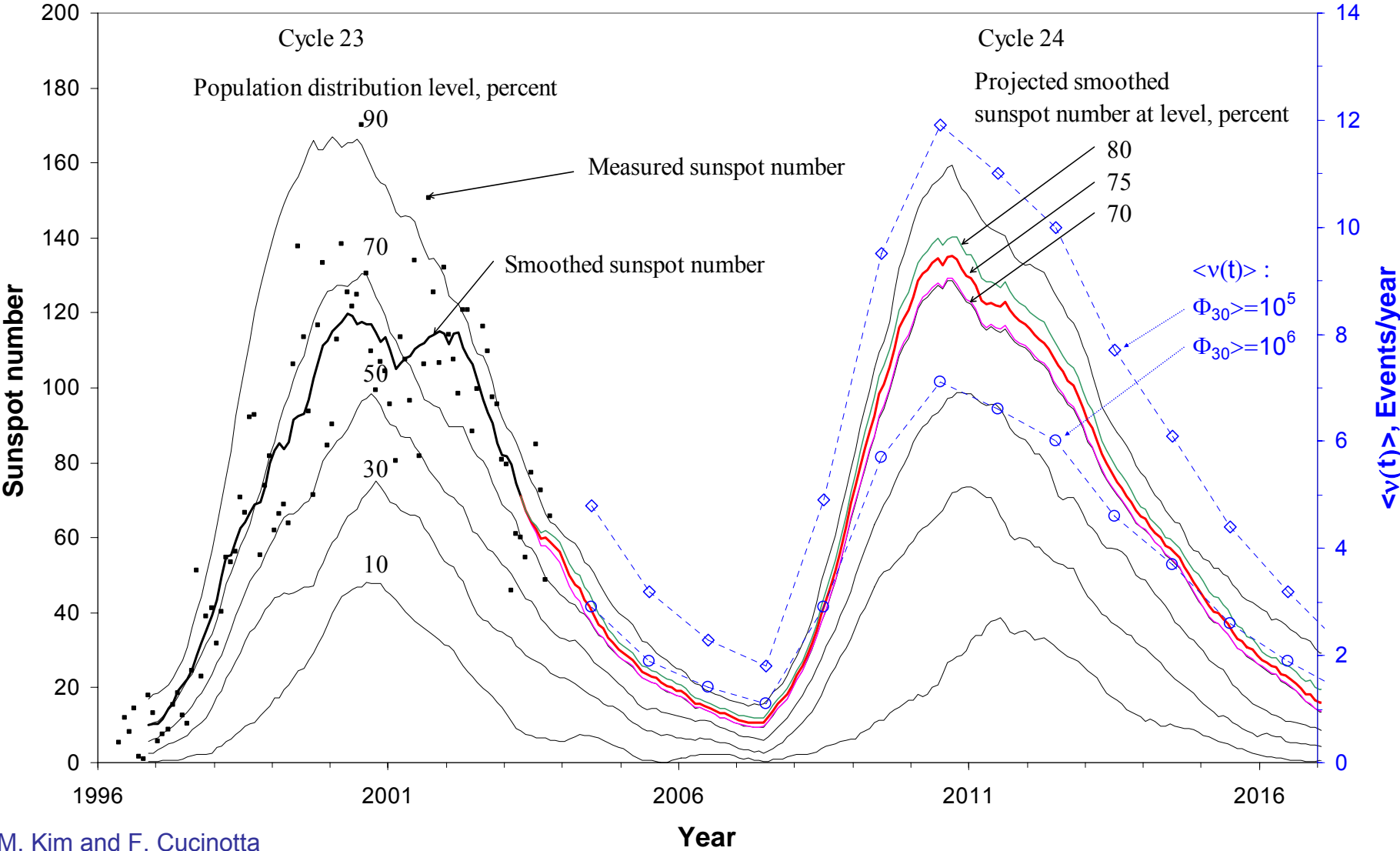


Fig. 1. The times of occurrence of >30 MeV solar proton events with fluence exceeding $1.0 \times 10^9/\text{cm}^2$, and the annual international sunspot numbers.

Recent Era (1550-2000)
McCracken et al.

Projections of Cycles and Mean Occurrence Frequency of SPE



Large Solar Proton Events during Solar Cycles 19-23 with $\Phi_{30} > 10^9$ protons/cm²

Cycle	Onset Time	Φ_{30} , protons/cm ²	Time to Peak Flux, hr	Comment
19	11/12/1960	9.00 x 10⁹	14	(a), note 1
20	8/2/1972	5.00 x 10⁹	69	(a)
22	10/19/1989 13:05	4.23 x 10 ⁹	26.9	(b)
23	7/14/2000 10:45	3.74 x 10 ⁹	25.8	(b)
23	10/26/2003 18:25	3.25 x 10⁹	4.2	(b), note 2
23	11/4/2001 17:05	2.92 x 10 ⁹	33.2	(b)
19	7/10/1959	2.30 x 10 ⁹	note 3	(a)
23	11/8/2000 23:50	2.27 x 10 ⁹	16.1	(b)
22	3/23/1991 8:20	1.74 x 10 ⁹	19.5	(b)
22	8/12/1989 16:00	1.51 x 10 ⁹	15.2	(b)
22	9/29/1989 12:05	1.35 x 10 ⁹	14.1	(b)
23	1/16/2005 2:10	1.04 x 10⁹	39.7	(b)
19	2/23/1956	1.00 x 10⁹	note 4	(a)

(a) Φ_{30} for solar cycle 19-21: data taken from Shea and Smart.

(b) Φ_{30} for solar cycle 22 and 23: calculated using corrected 5-min average proton flux of GOES measurements.

Note 1: There are large differences in the estimate of fluence in November 1960 SPE.

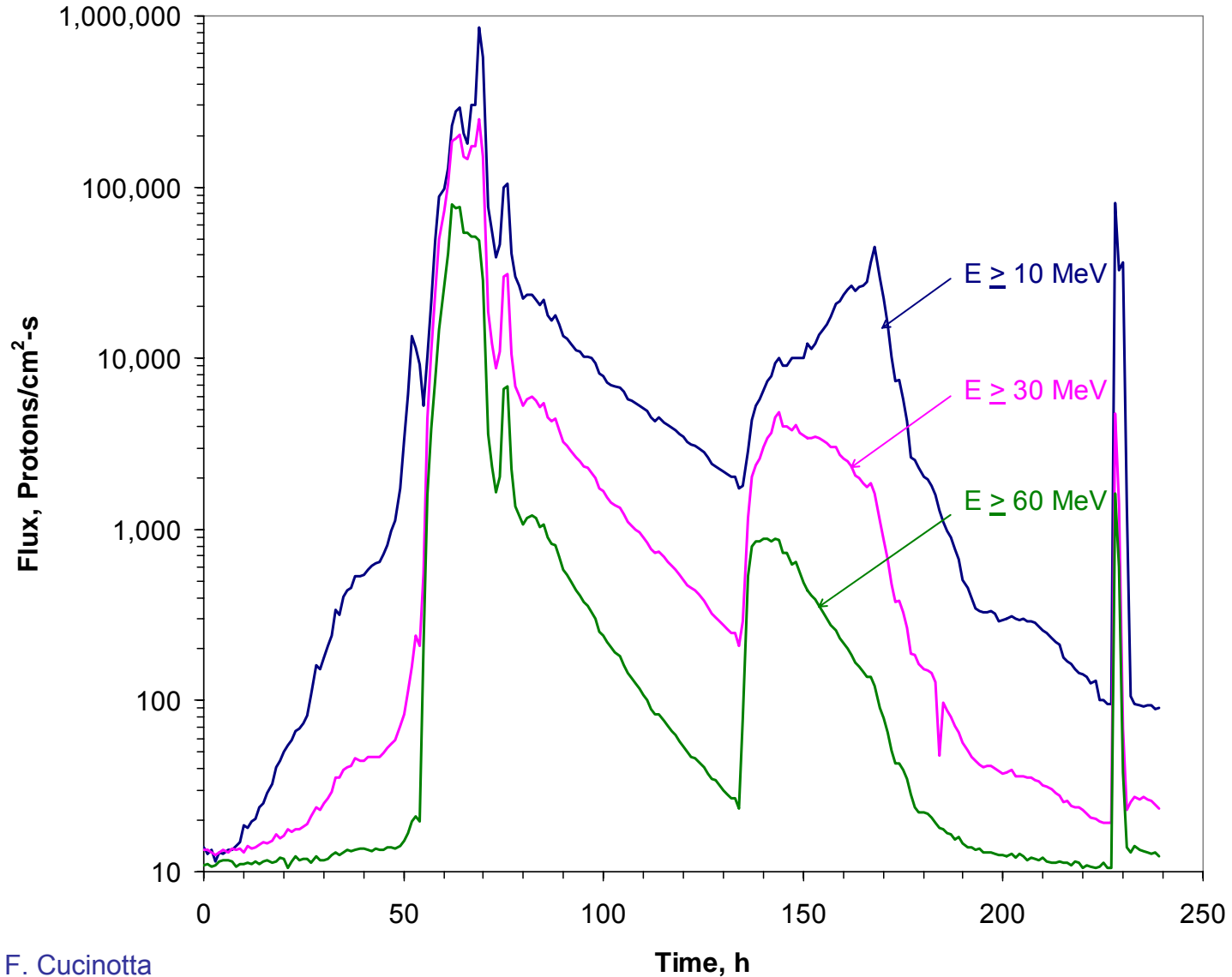
Data given by others are significantly smaller than this value (1.3 x 10⁹ protons/cm² by Freier and Webber).

Note 2: Φ_{30} for the combined 3 major peaks occurred during 10/26-11/6/2003: 3.42 x 10⁹ protons/cm².

Note 3: one day later on July 11, 1959

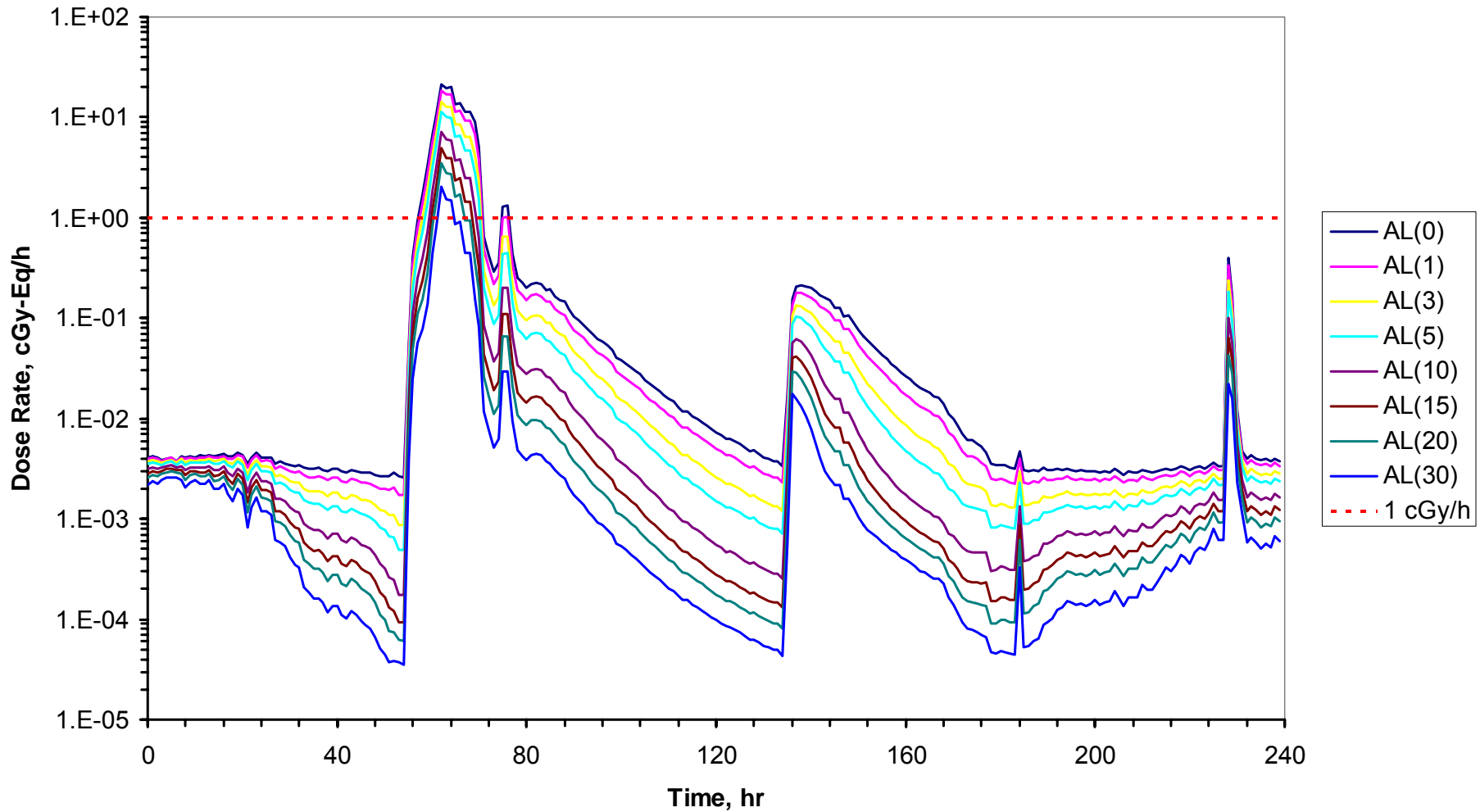
Note 4: in the same day on February 23, 1956

Proton Integral Flux August 2-11, 1972 SPE



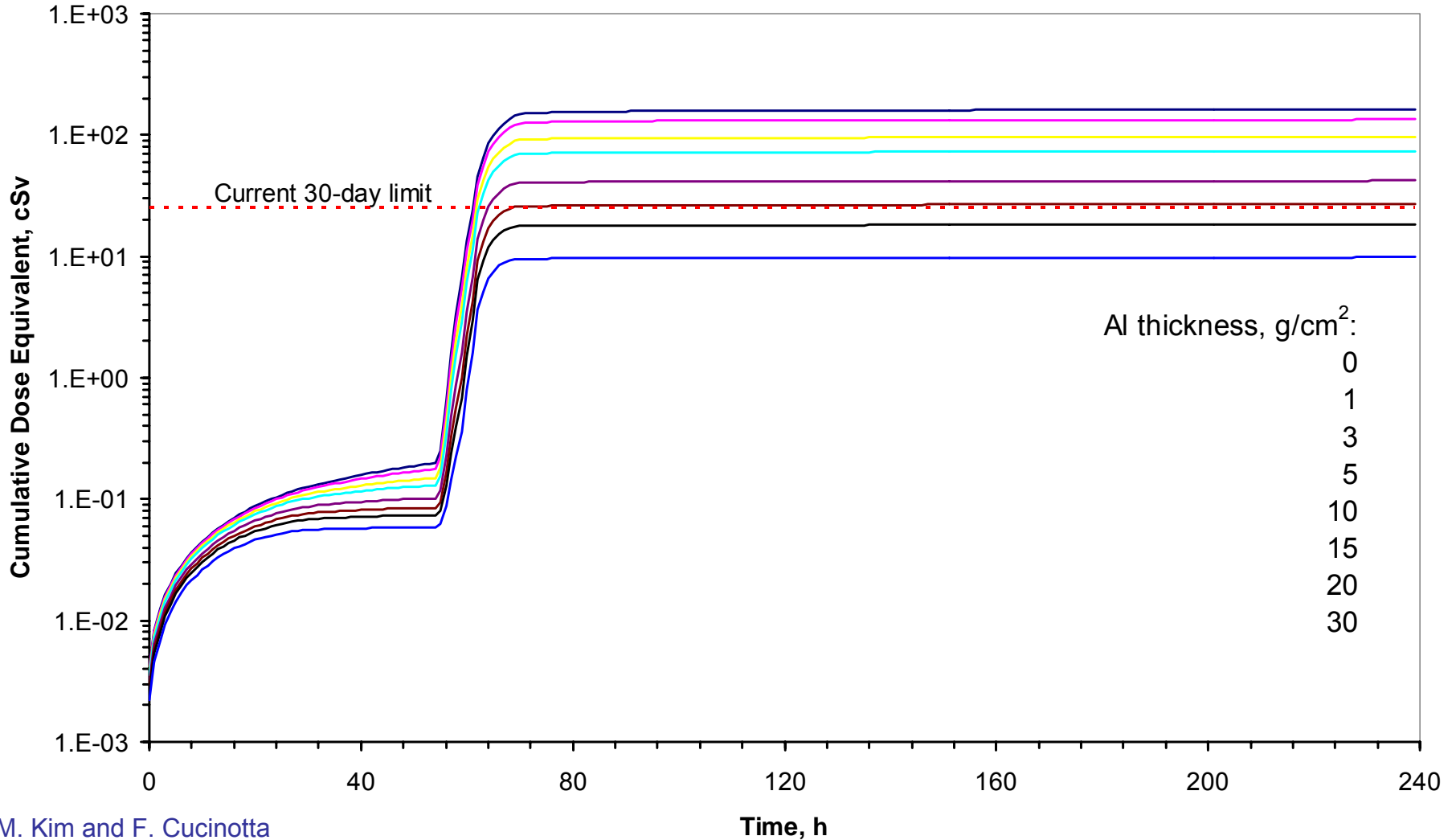
BFO Dose Rate

August 2-11, 1972 SPE

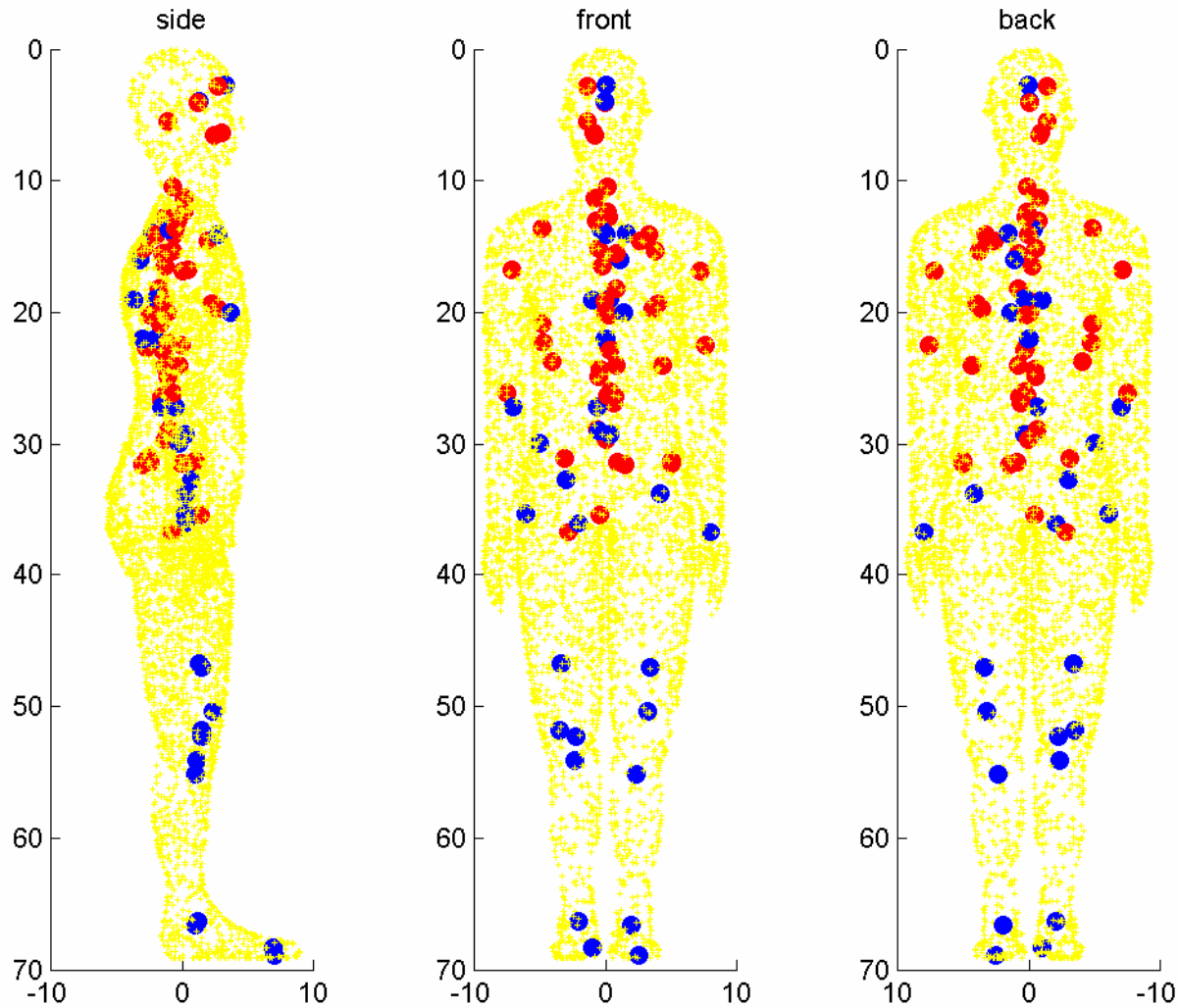


BFO Cumulative Dose

August 2-11, 1972 SPE

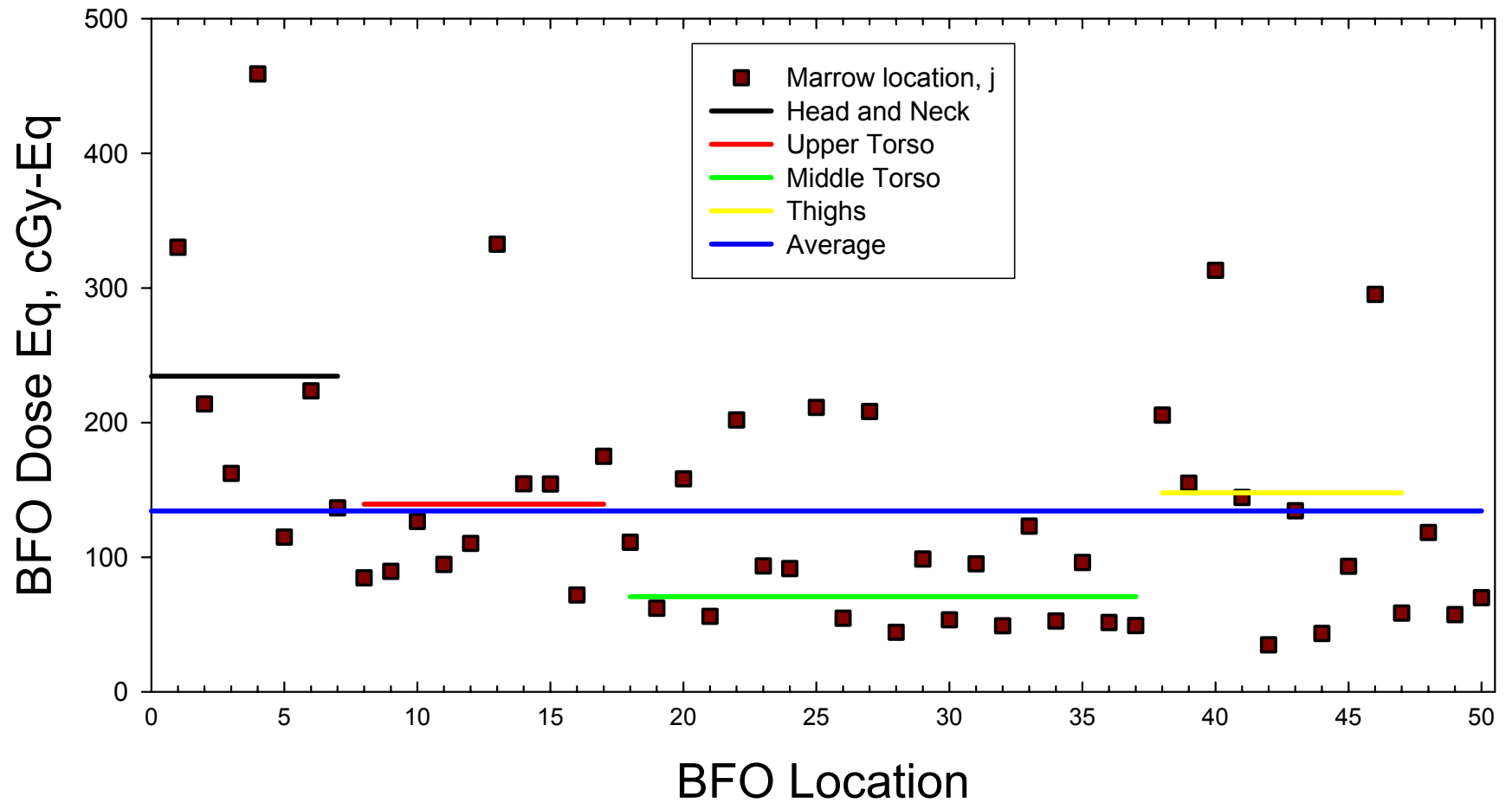


82 BFO Locations

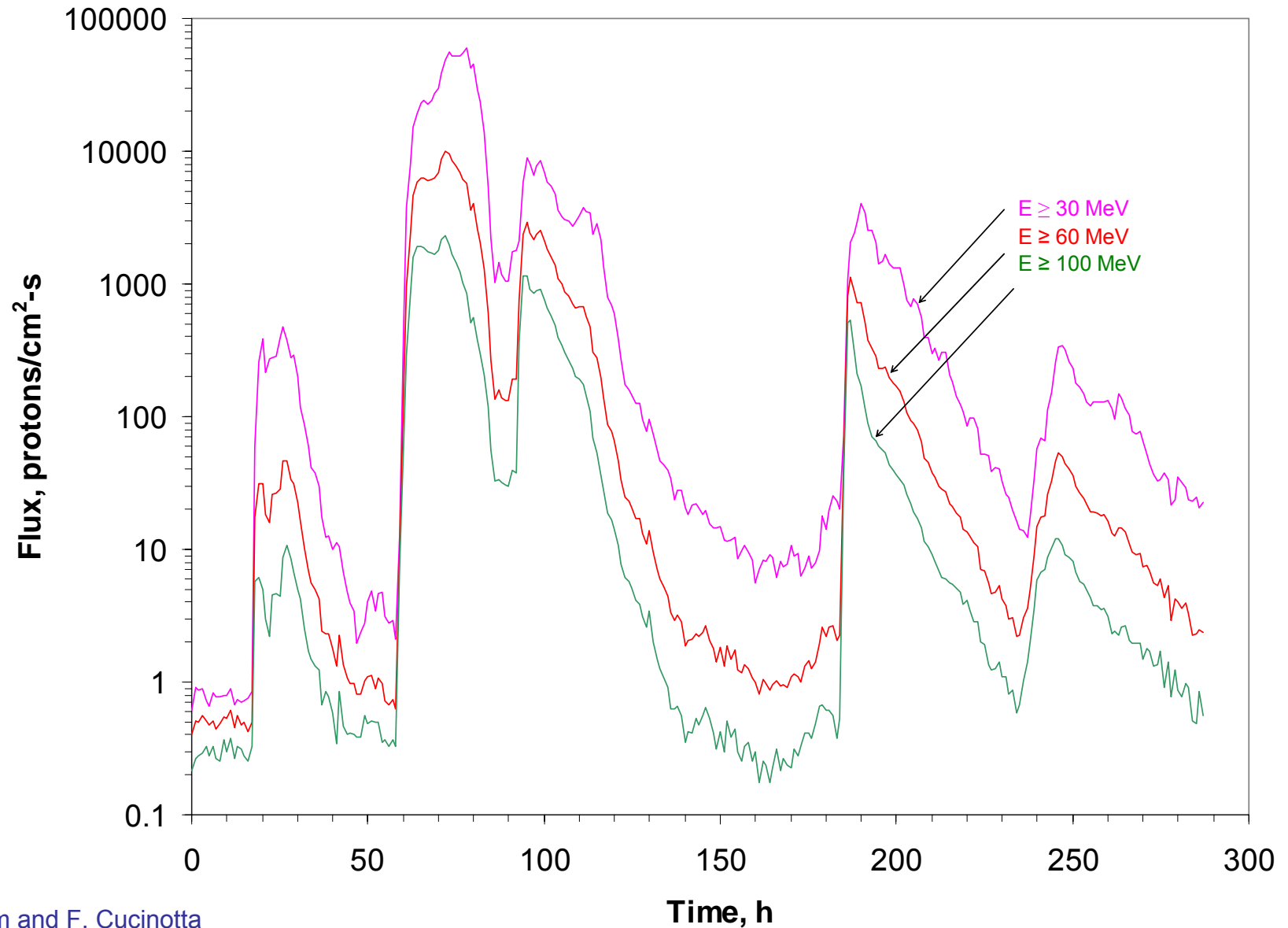


- 32 location set
- 50 location set

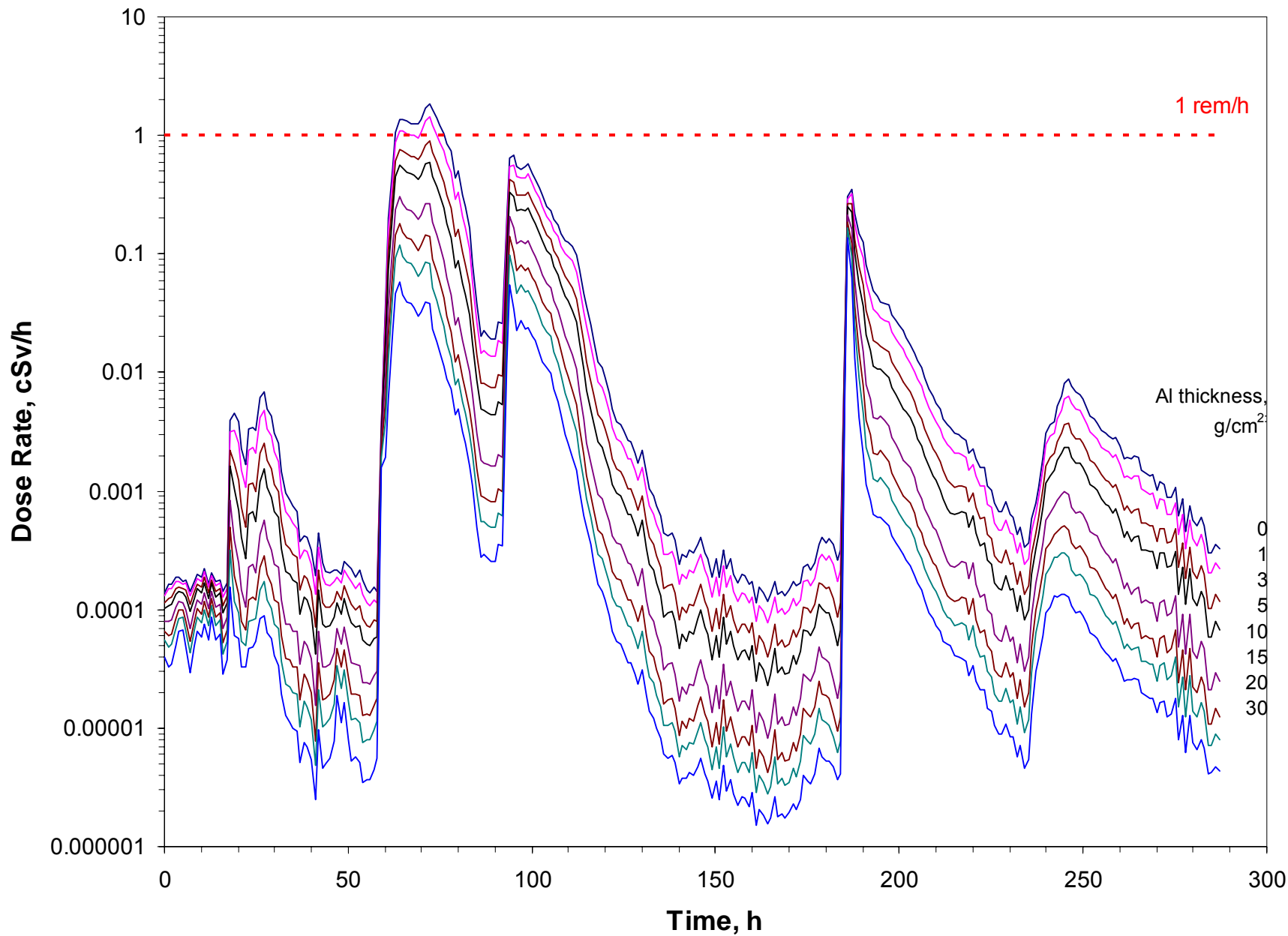
August 1972 SPE (1 g/cm² Al shield)



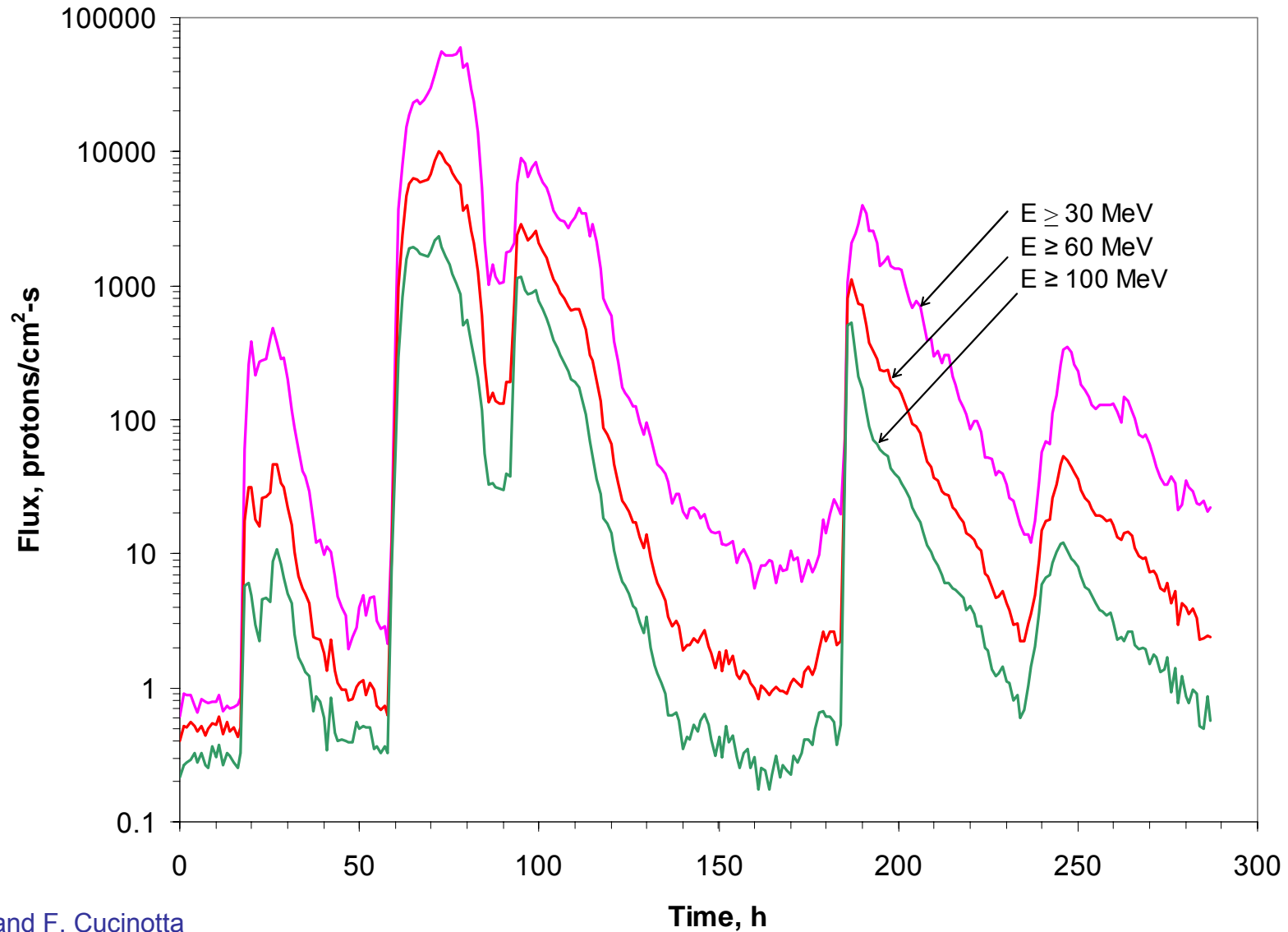
Hourly-Averaged Proton Integral Flux during Oct 26 - Nov 6, 2003 SPE



BFO Dose Rate during Oct 26 - Nov 6, 2003 SPE

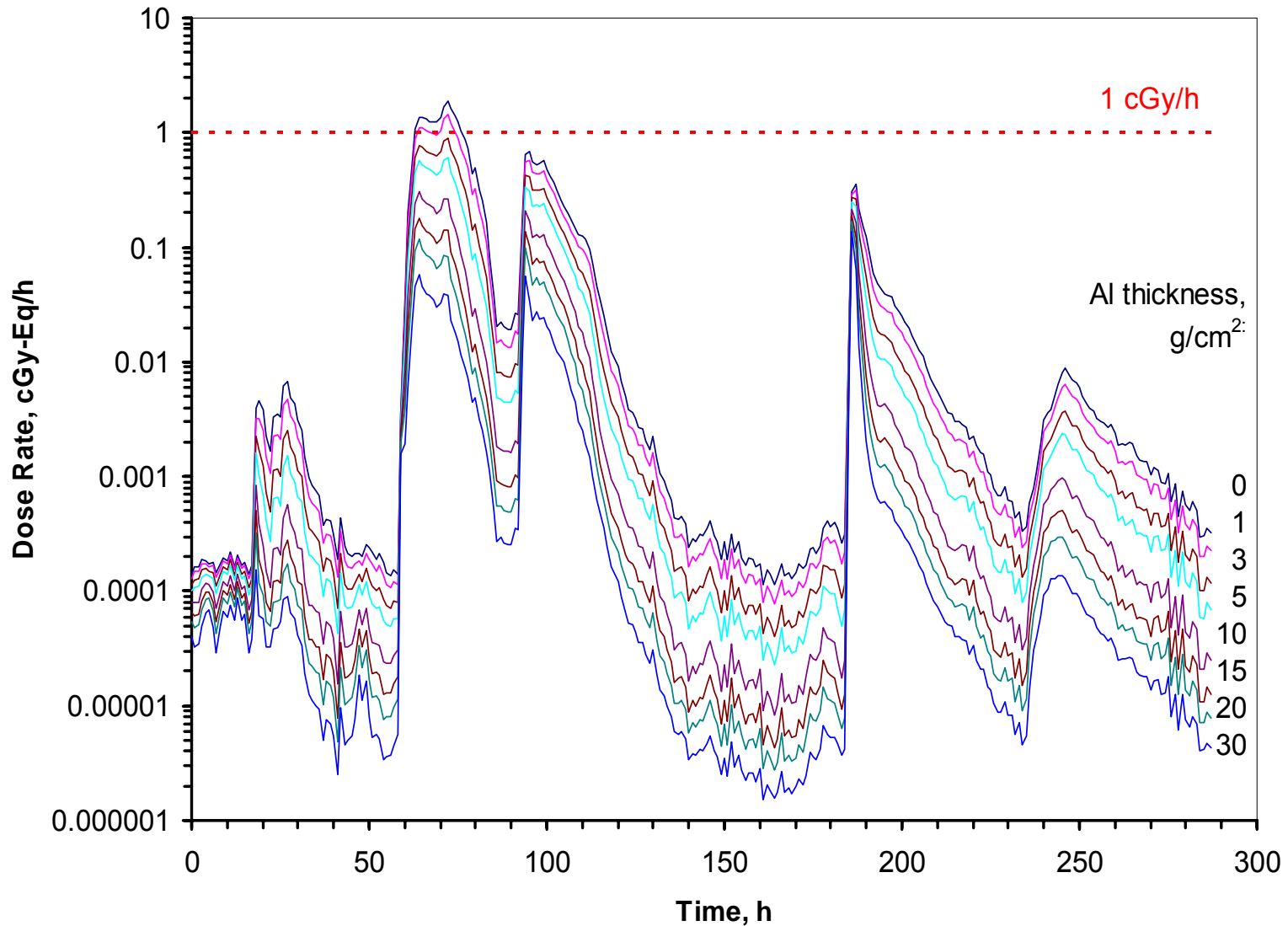


Hourly-Averaged Proton Integral Flux Oct 26 - Nov 6, 2003 SPE



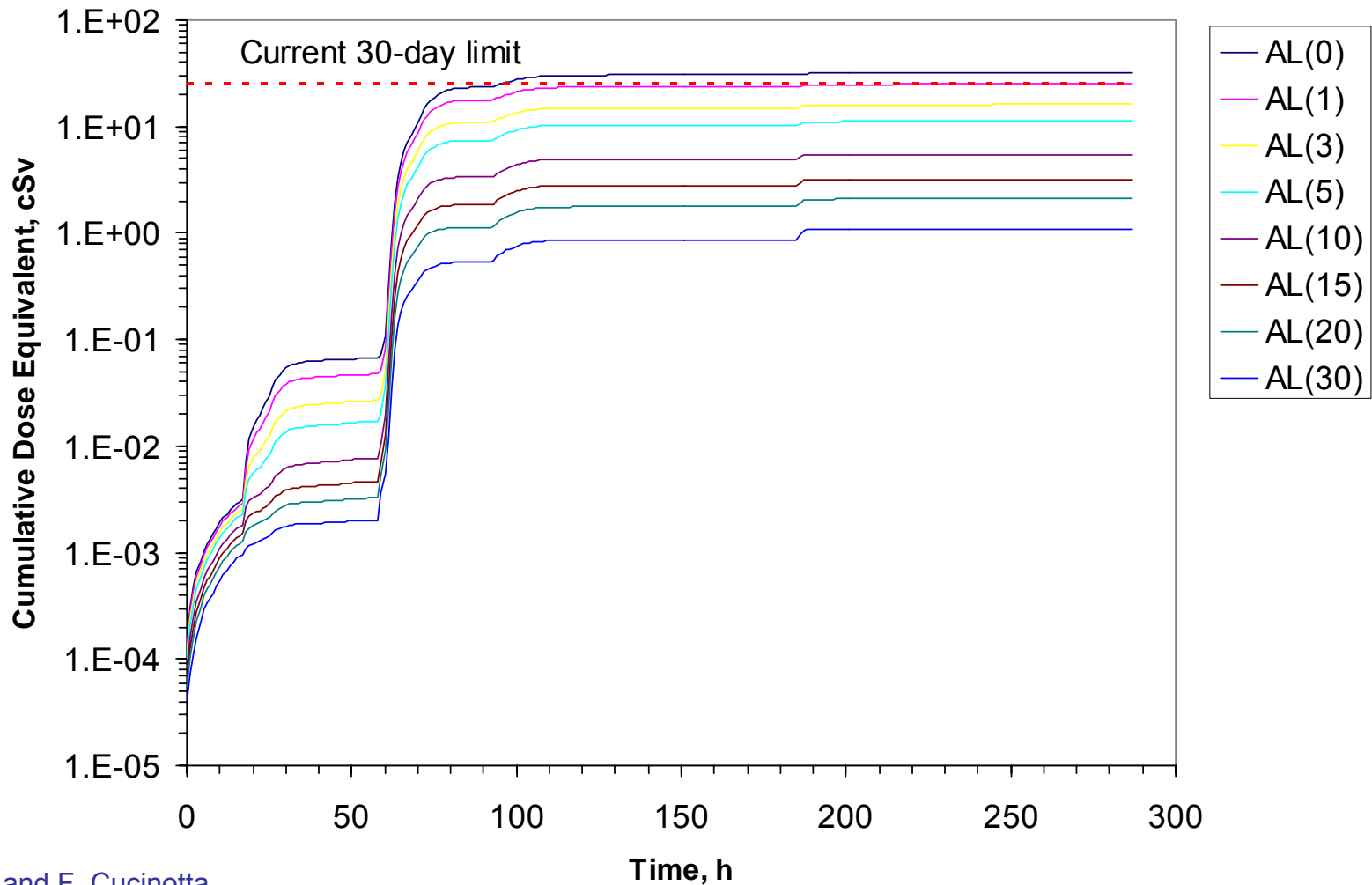
BFO Dose Rate

October 26- November 6, 2003 SPE

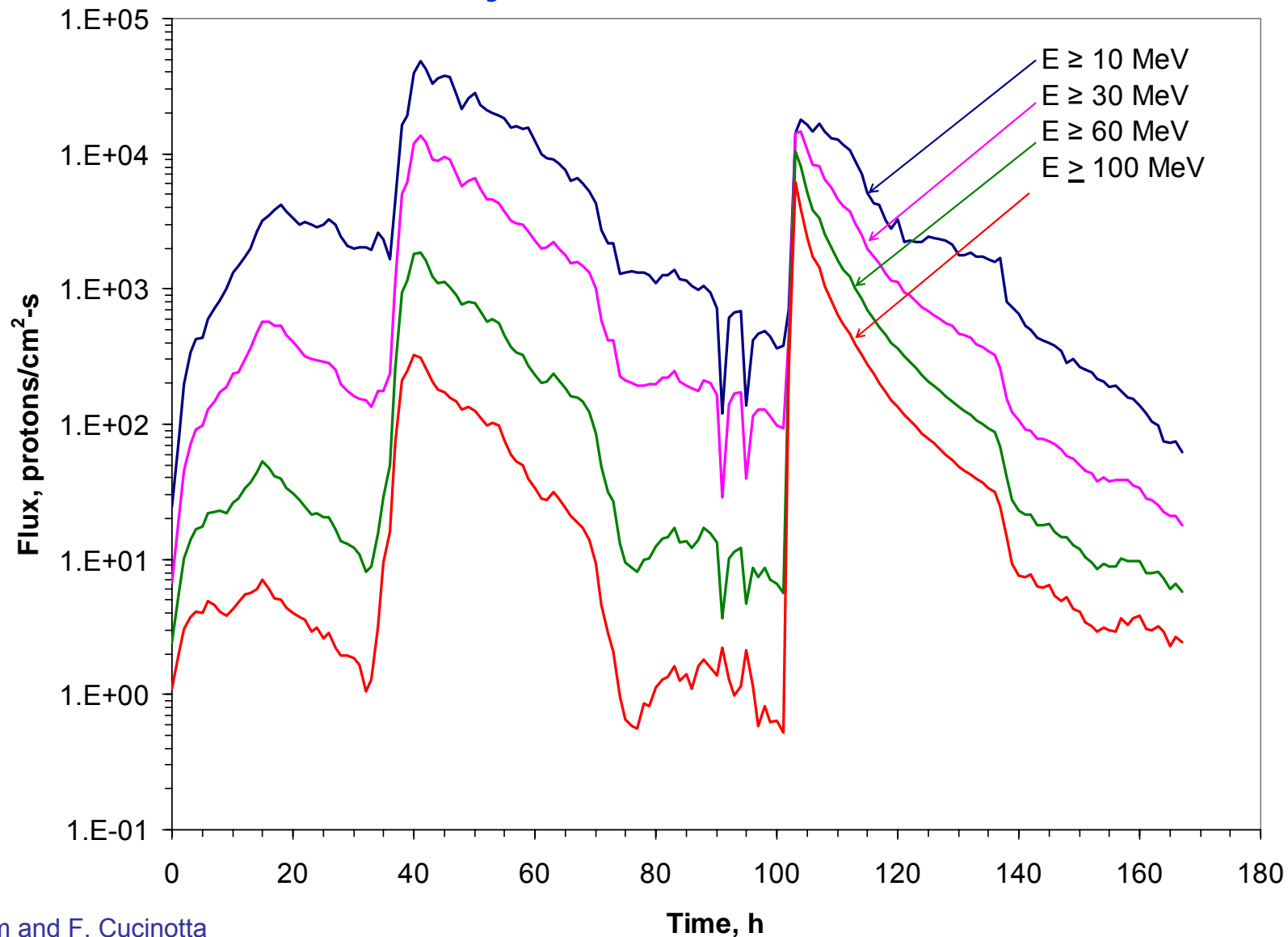


BFO Cumulative Dose

October 26- November 6, 2003 SPE

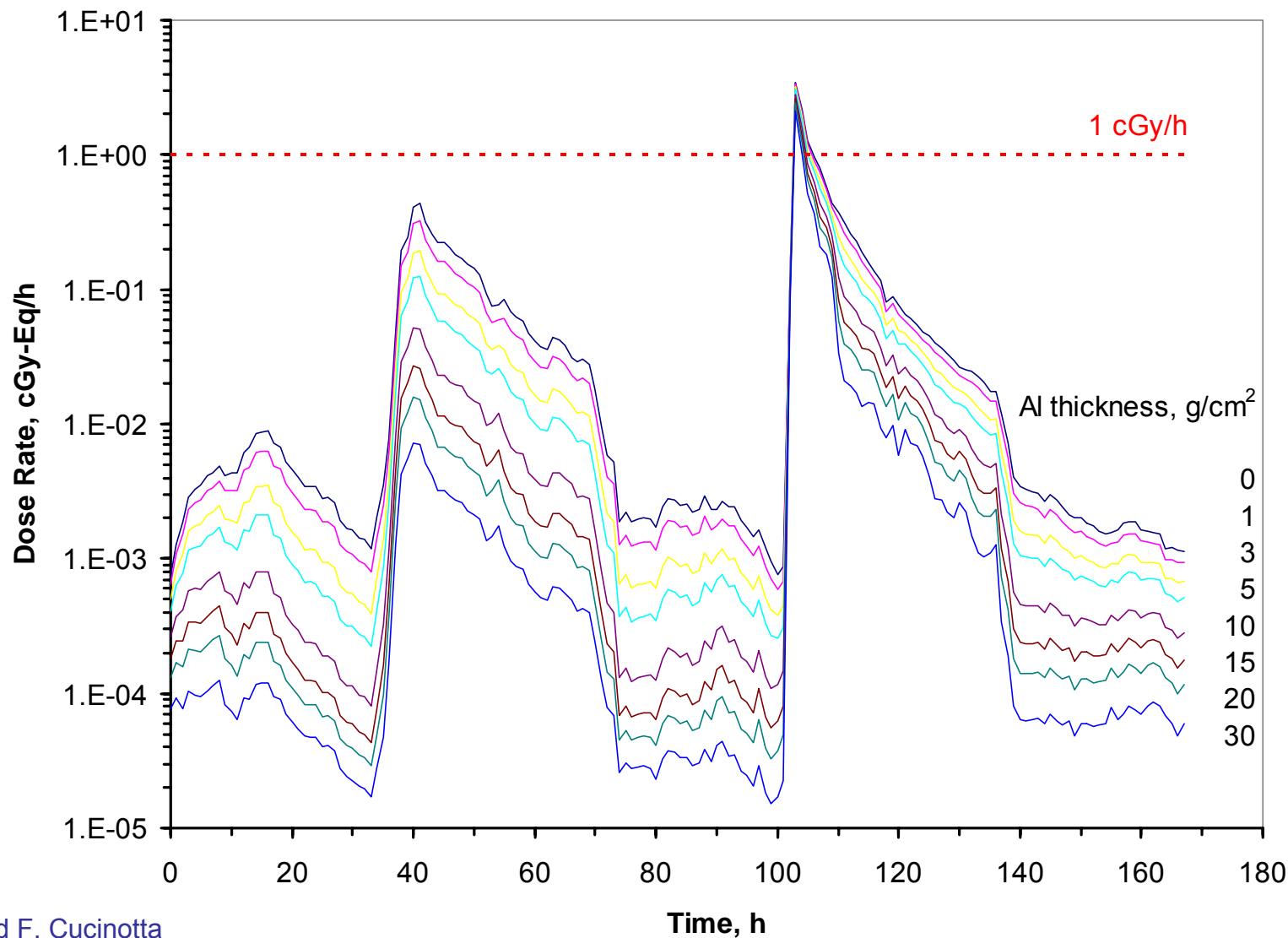


Hourly-Averaged Proton Integral Flux January 16-22, 2005 SPE

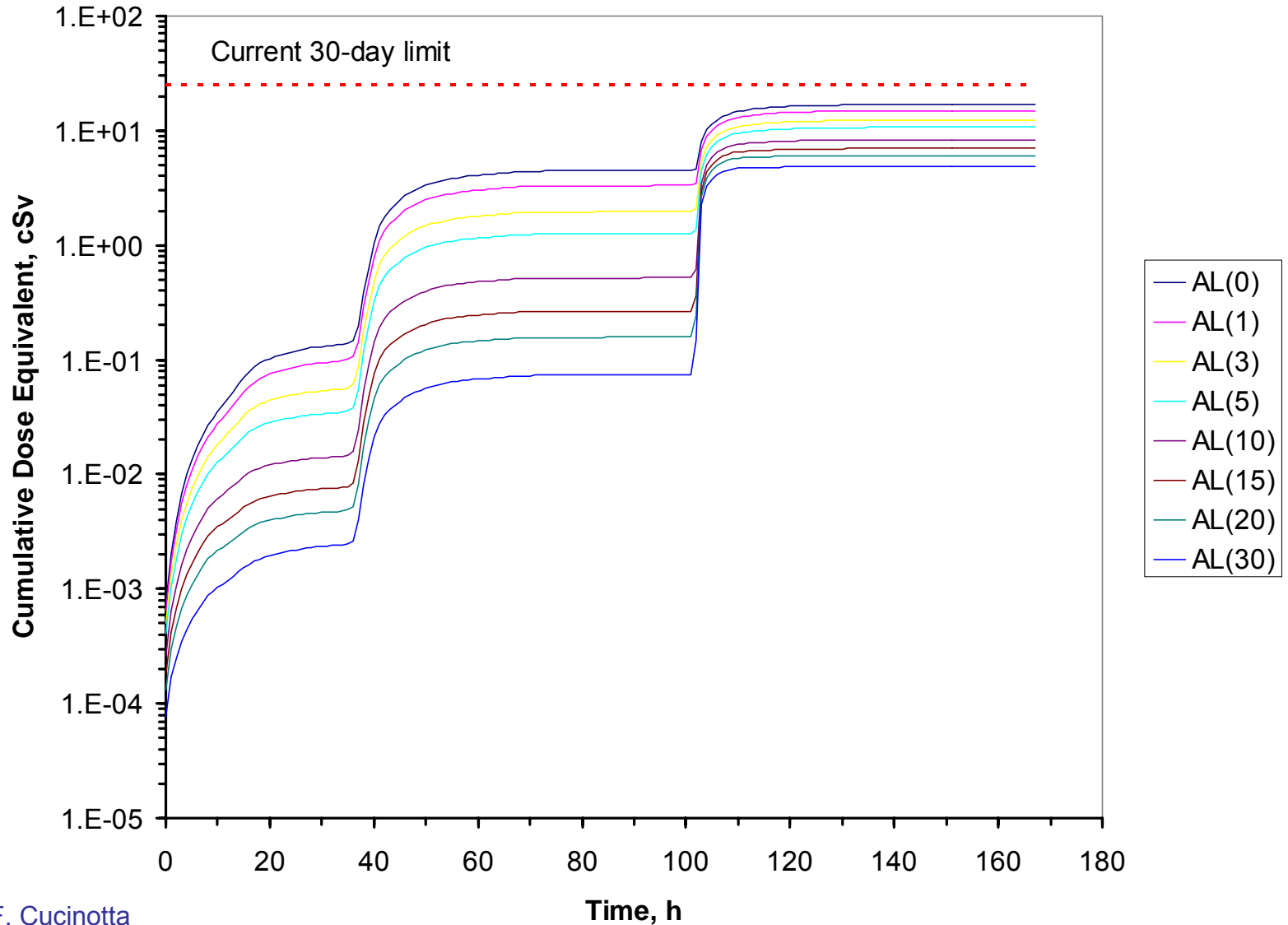


BFO Dose Rate

January 16-22, 2005 SPE



BFO Cumulative Dose January 16-22, 2005 SPE



Role of Dose-Rate and Shielding

- ❑ Shielding mitigates most SPE events
 - High-energy component (>100 MeV) often poorly known
- ❑ Proton biological damage is dependent on dose-rate
 - Effects increase above ~ 5 rad/hr

CAM ANALYSIS OF THE OCTOBER 1989 PROTON EVENT

EVA Today (Y/N?)
IVA (time to storm shelter)

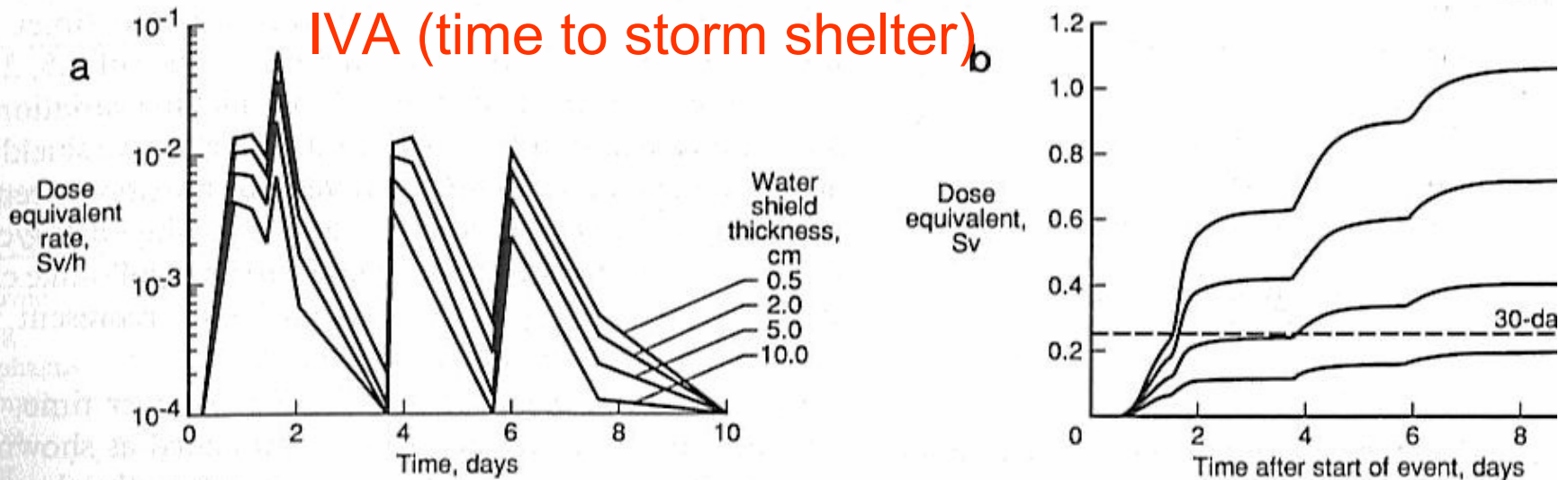
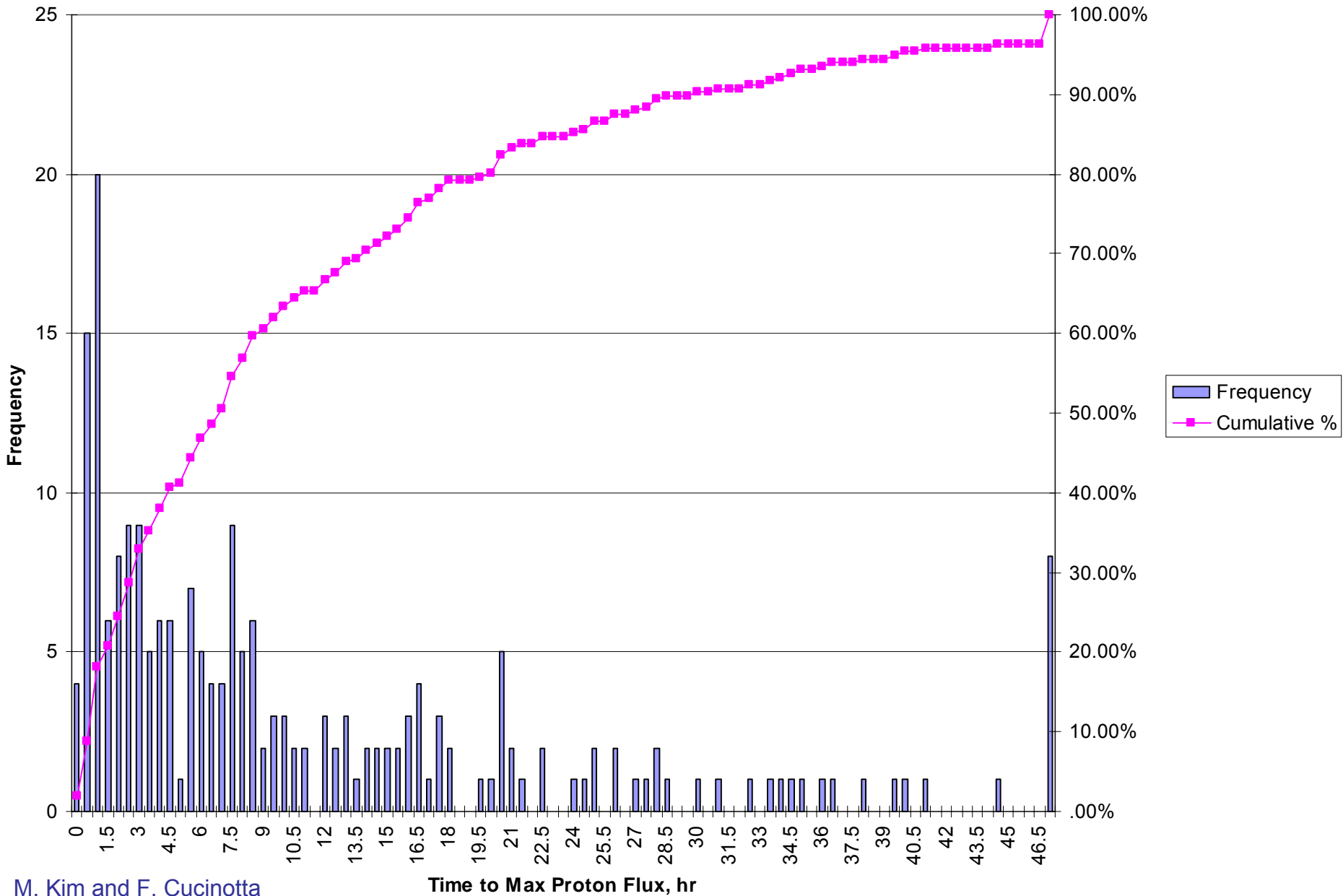
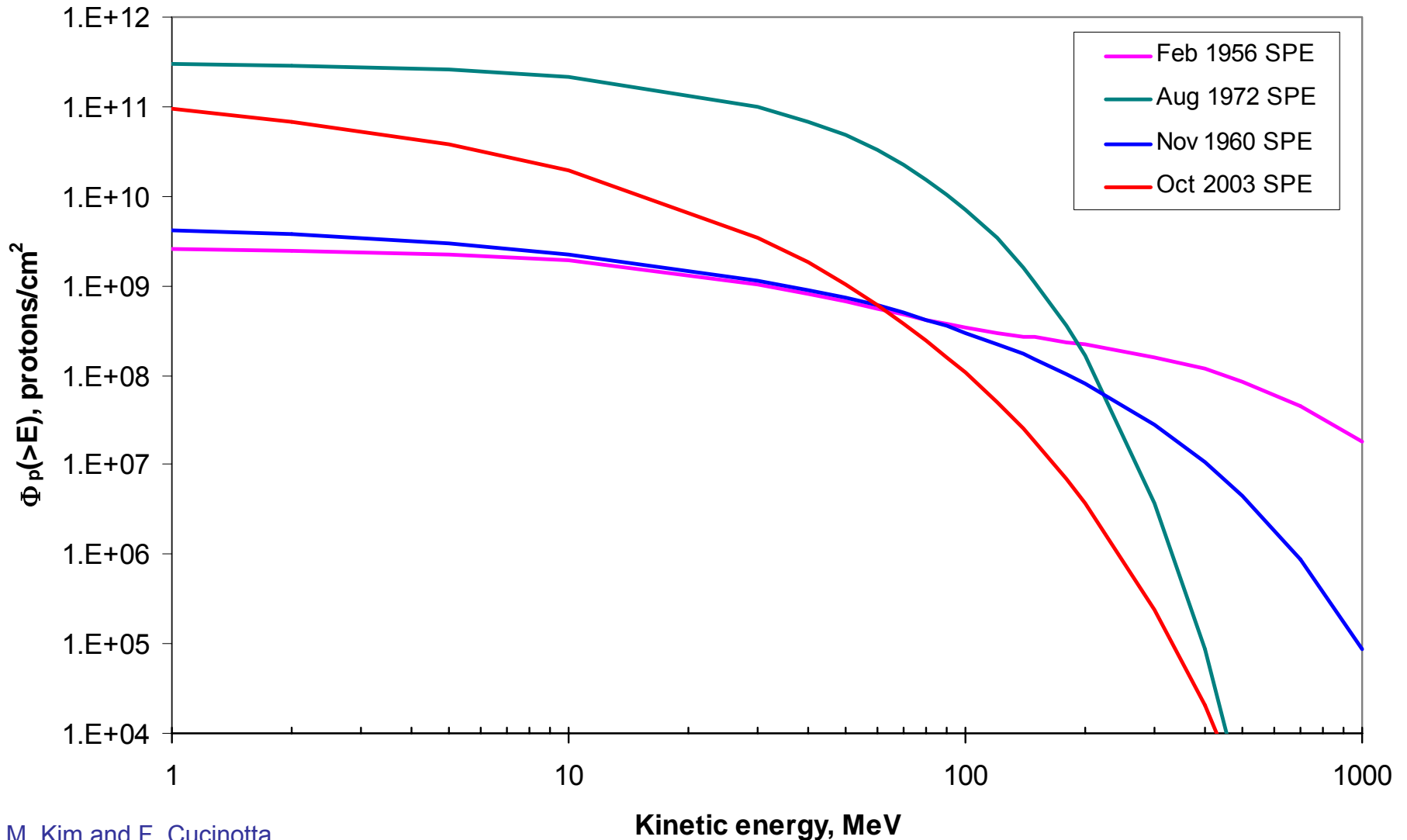


FIG. 9. Predicted BFO dose-rate variation and cumulative dose during the October 1989 event with 0.5, 2.0, 5.0, and 10.0 cm water shield thickness. (a) BFO dose equivalent rate; (b) cumulative BFO dose equivalent.

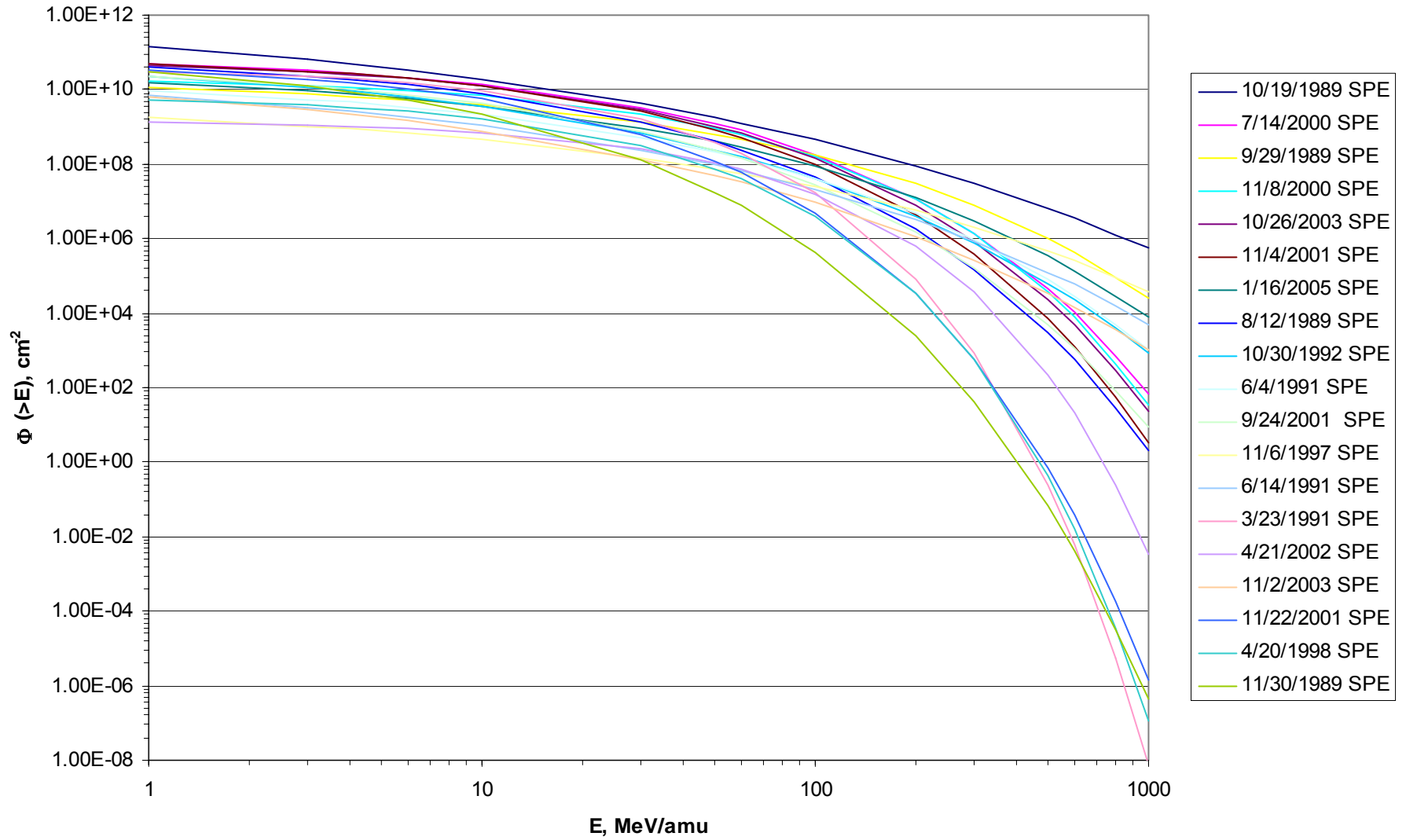
216 Particle Events (1976-2005)

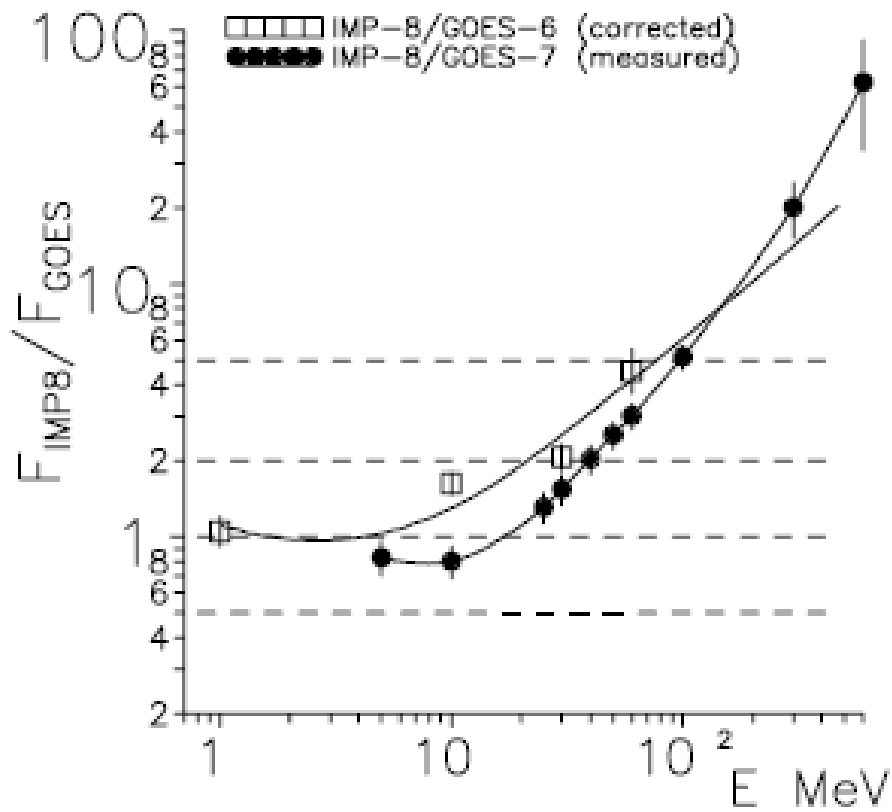


Large SPE Integral Fluence Spectra at 1AU



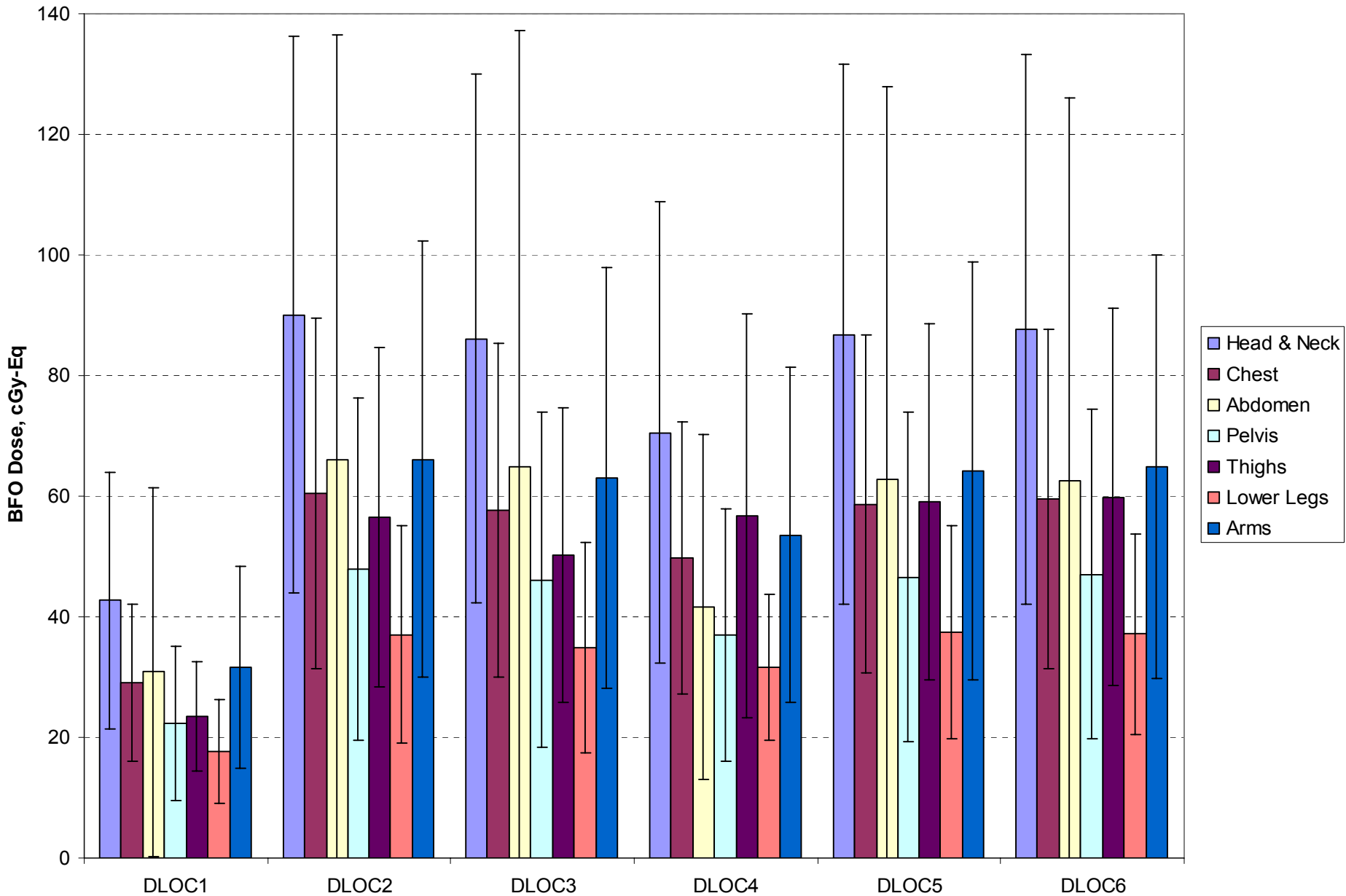
Big SPEs (Cycles 22-23)





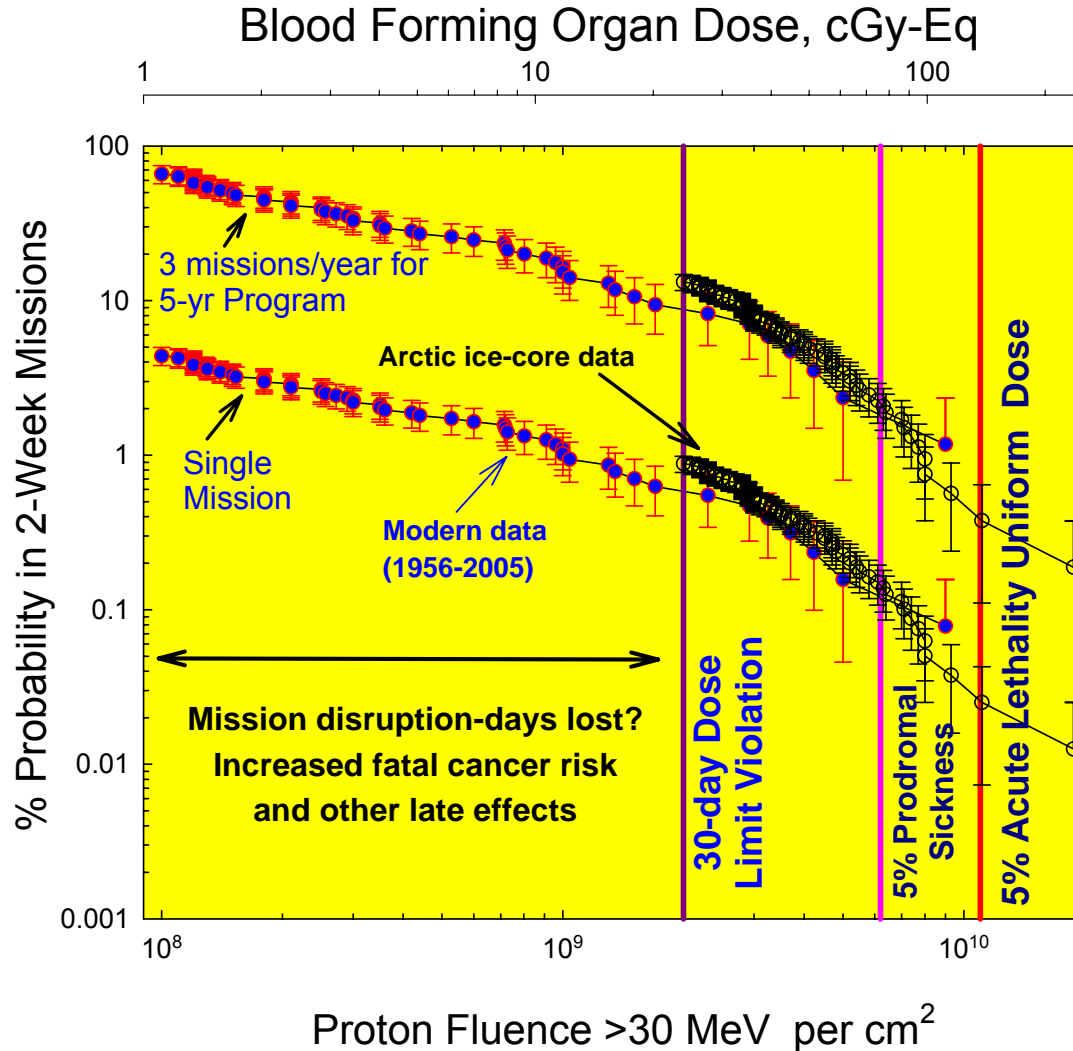
From Nymik- COSPAR

The Average BFO Doses of the Regions at 6 DLOCs from 1972 SPE



Was Carrington Event (1859) Lethal?

1859 event has largest $F > 30$ Flux known



Conditions under
Apollo Spacecraft
-type shielding

Ice-Core data
- $F_{>30\text{MeV}}$ flux from
1450-1990

Large SPE's

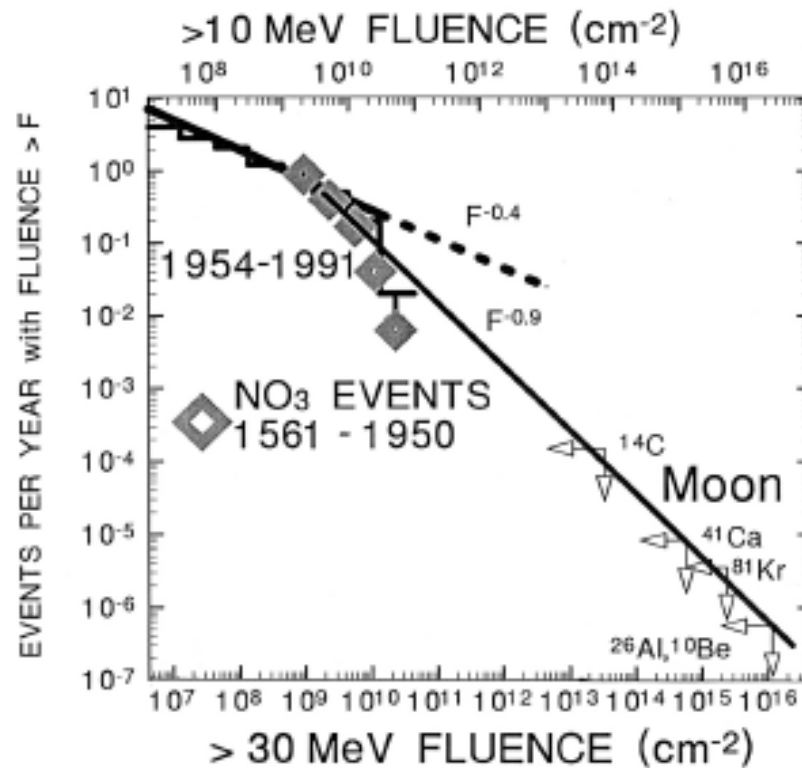
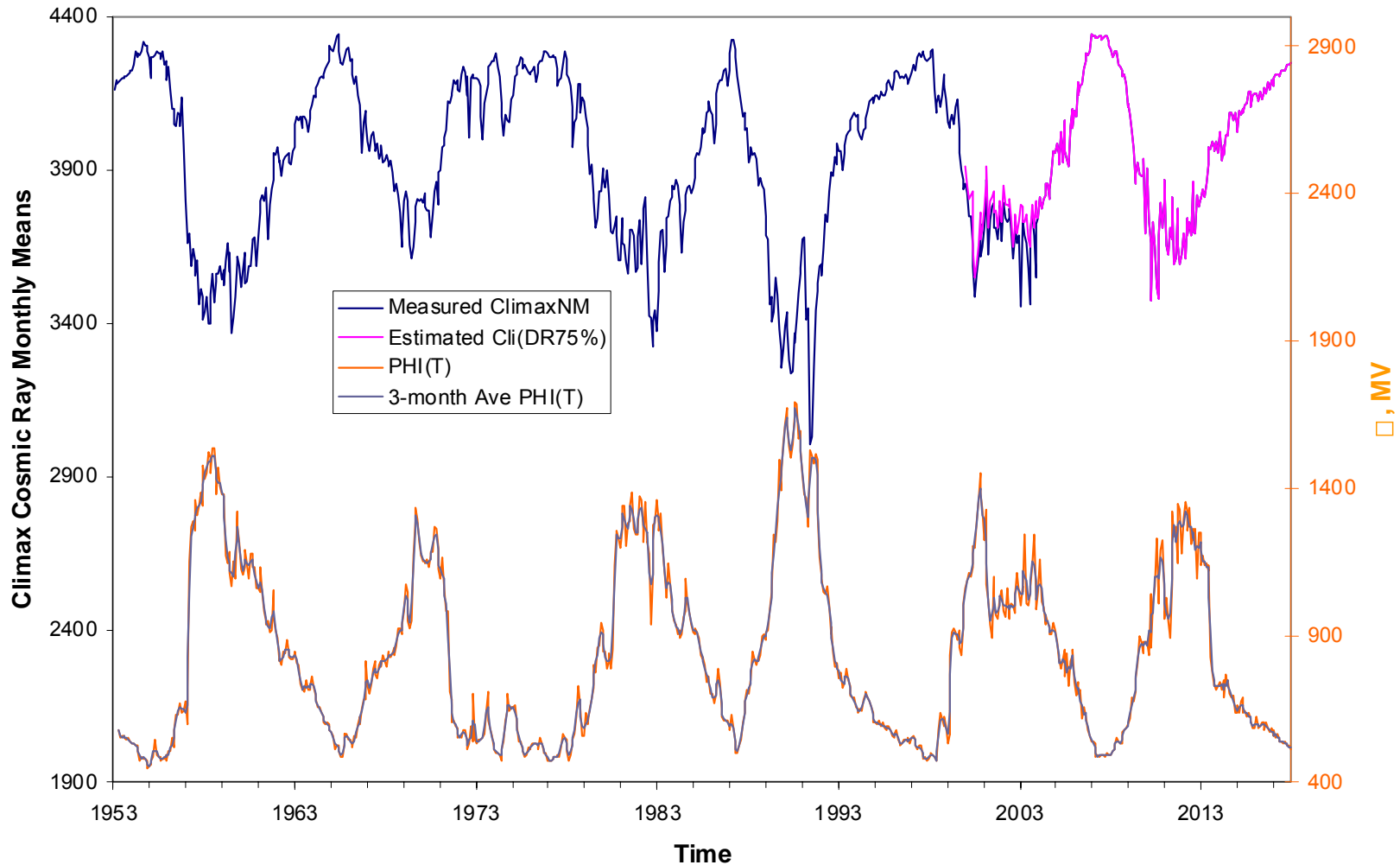


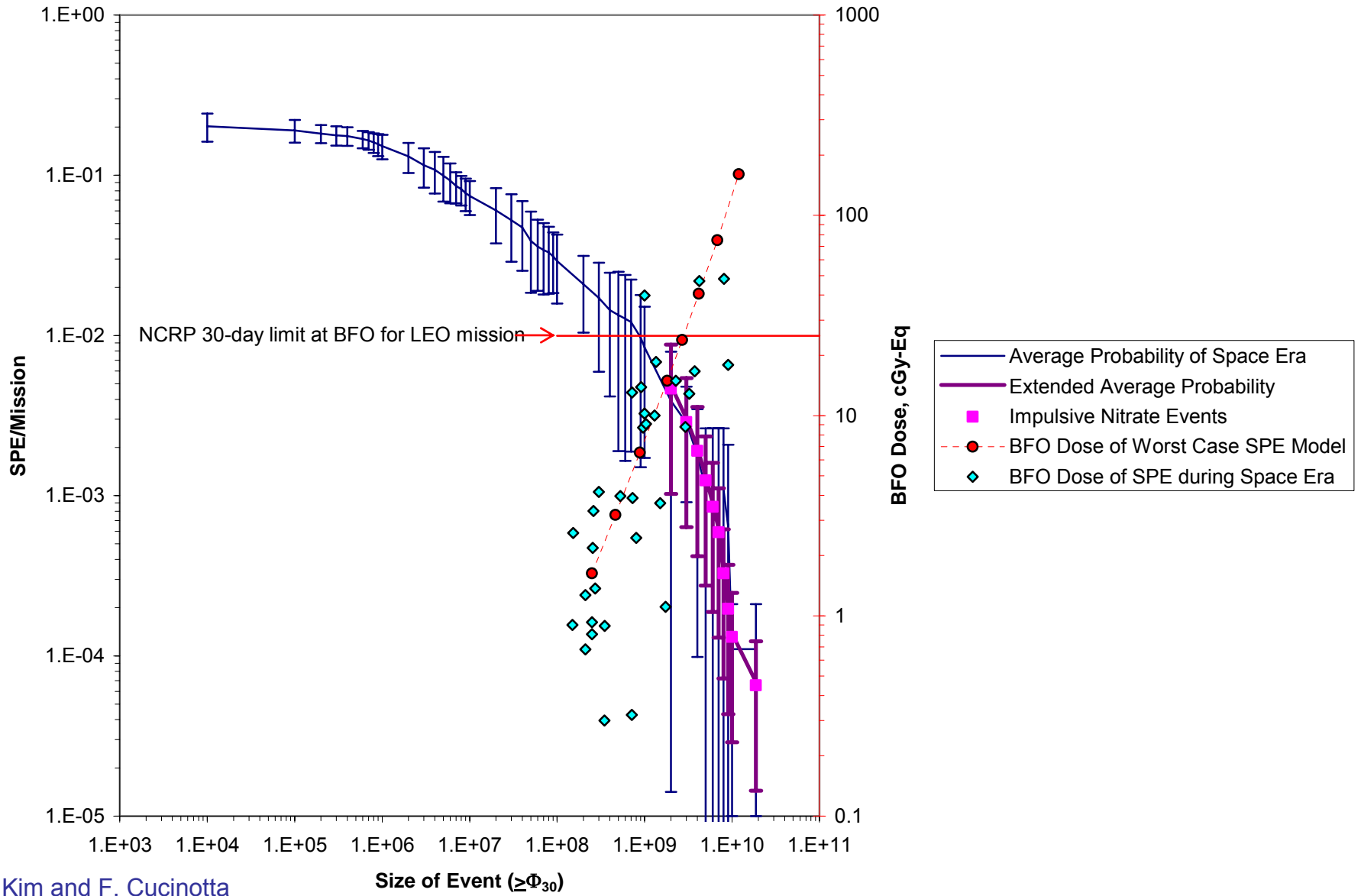
Fig.4. Cumulative probabilities of the SPEs observed by satellite, and derived from the nitrate data. The diamond shaped symbols refer to the nitrate data; the histogram and lines to the satellite data and fluence limits derived from cosmogenic isotopes in moon rocks.

Long-Term Forecasting

Climax Neutron Monitor and Deceleration Potential (Φ)

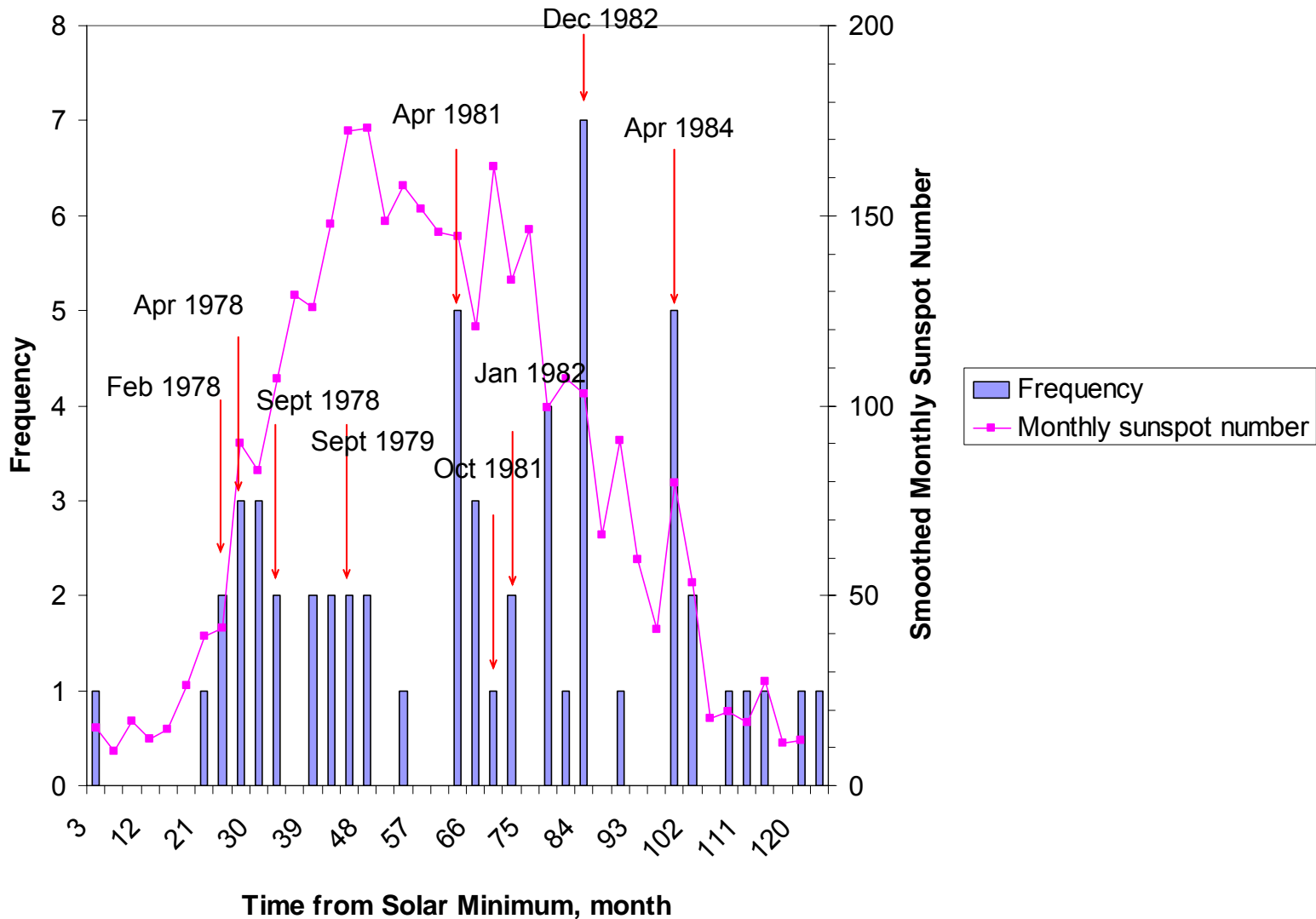


SPE Probability in 1-Week Mission and BFO Exposure Level inside a Typical Equipment Room in Free Space



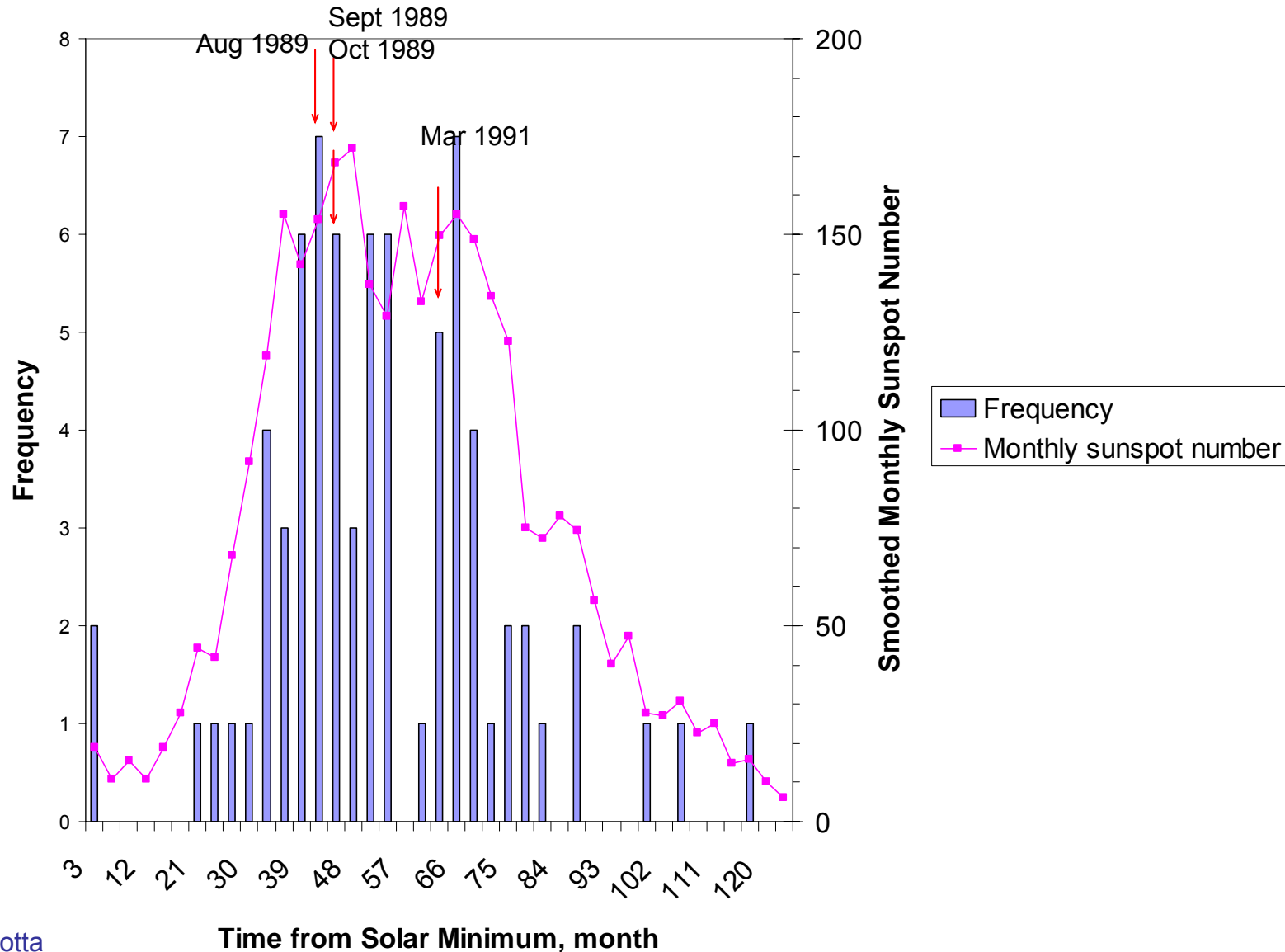
Frequency of SPE Occurrence in 3-Month Periods

$\Phi_{30} > 10^8$ protons/cm² for Solar Cycle 21.



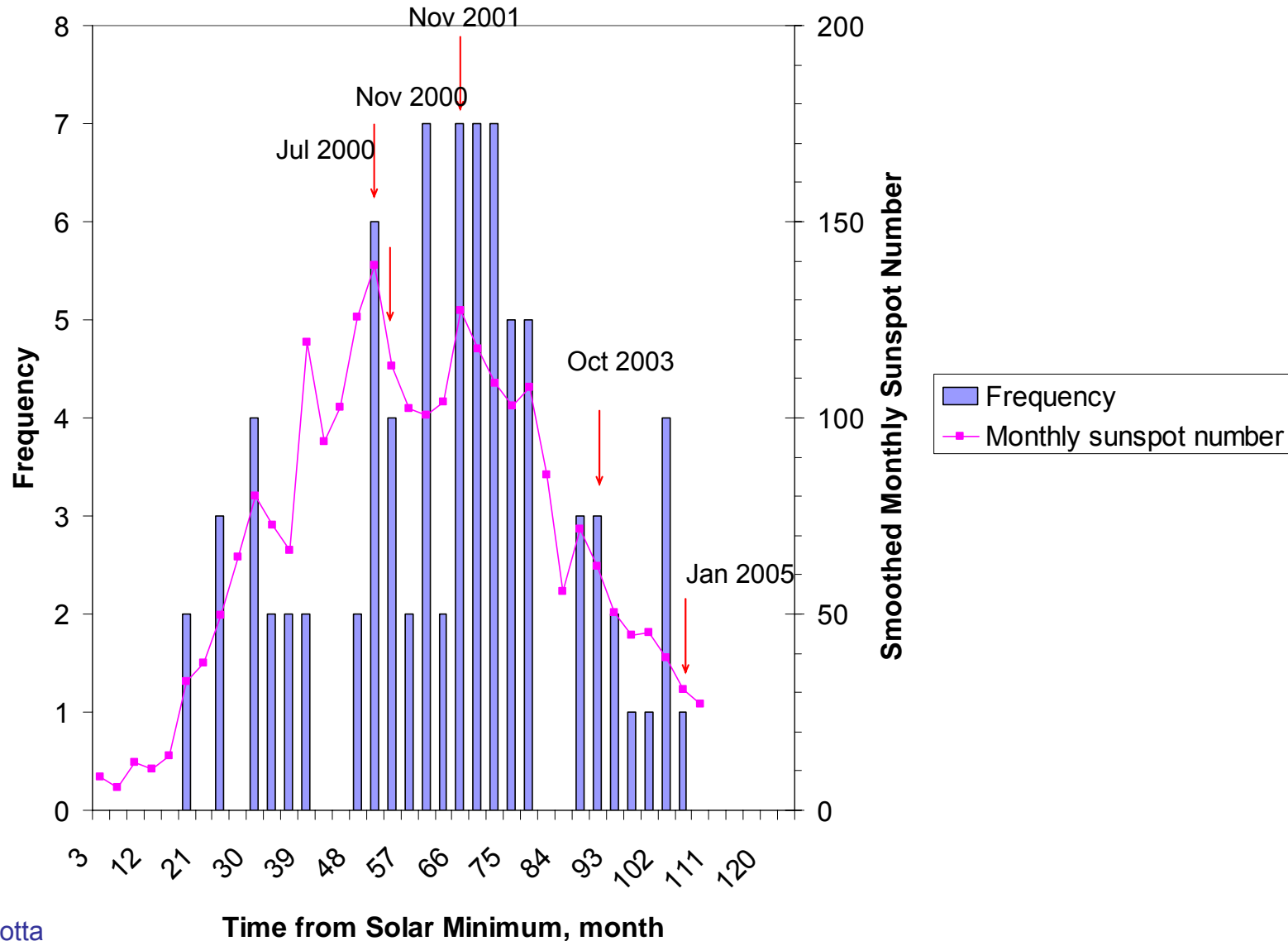
Frequency of SPE Occurrence in 3-Month Periods

$\Phi_{30} > 10^9$ protons/cm² for Solar Cycle 22.



Frequency of SPE Occurrence in 3-Month Periods

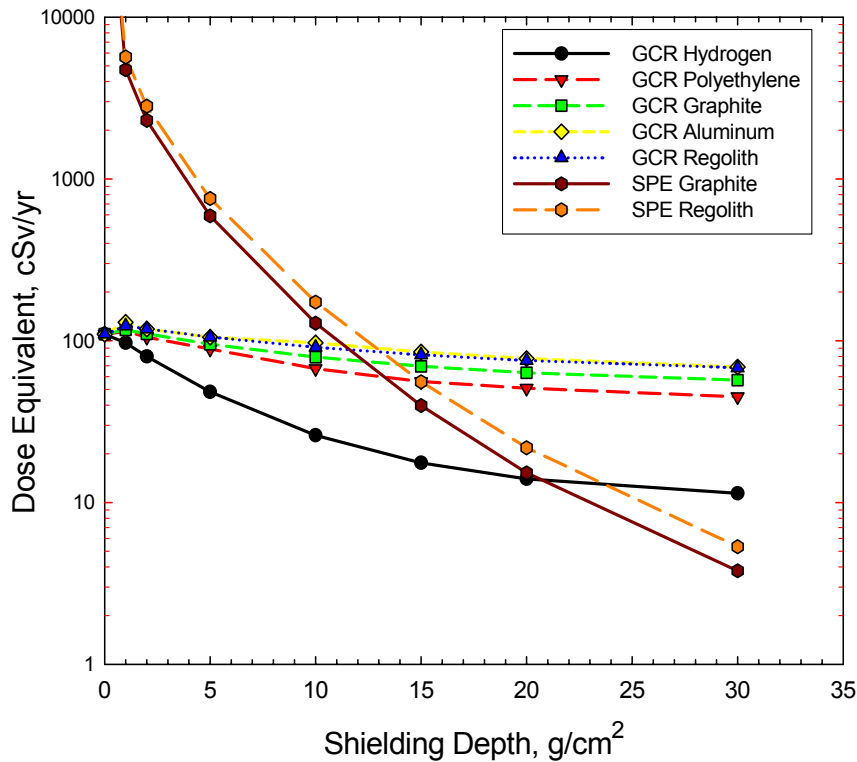
$\Phi_{30} > 10^9$ protons/cm² for Solar Cycle 23.



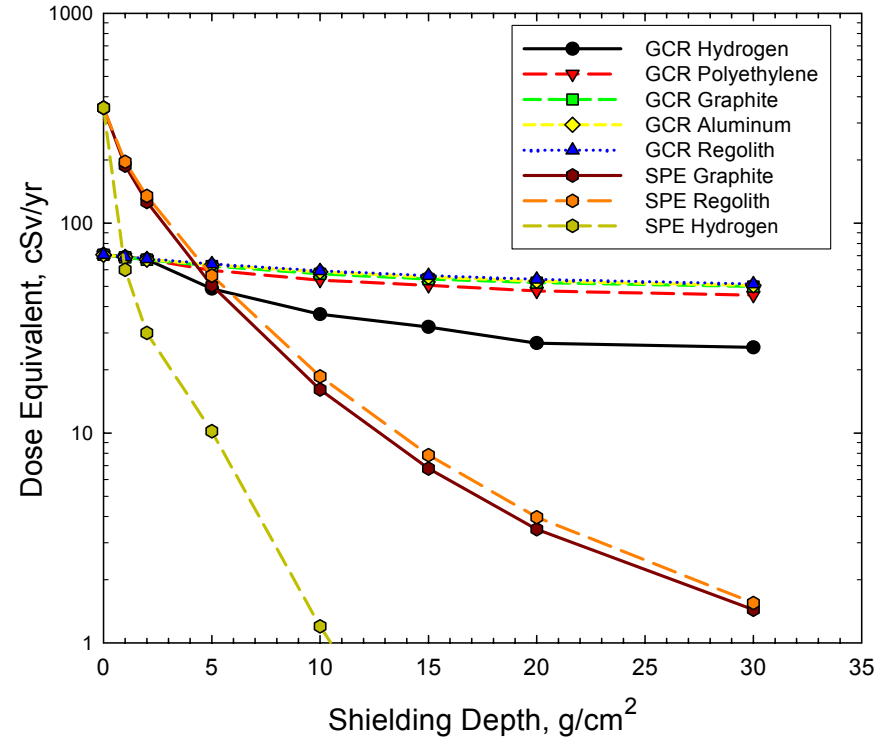
GCR and SPE Dose: Materials & Tissue

GCR higher energy >> secondary radiation

No Tissue Shielding

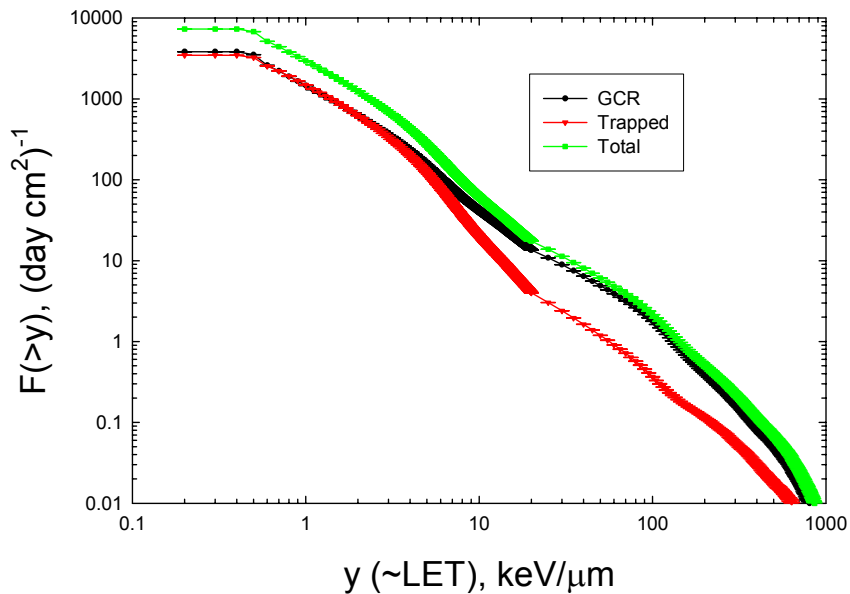


With Tissue Shielding

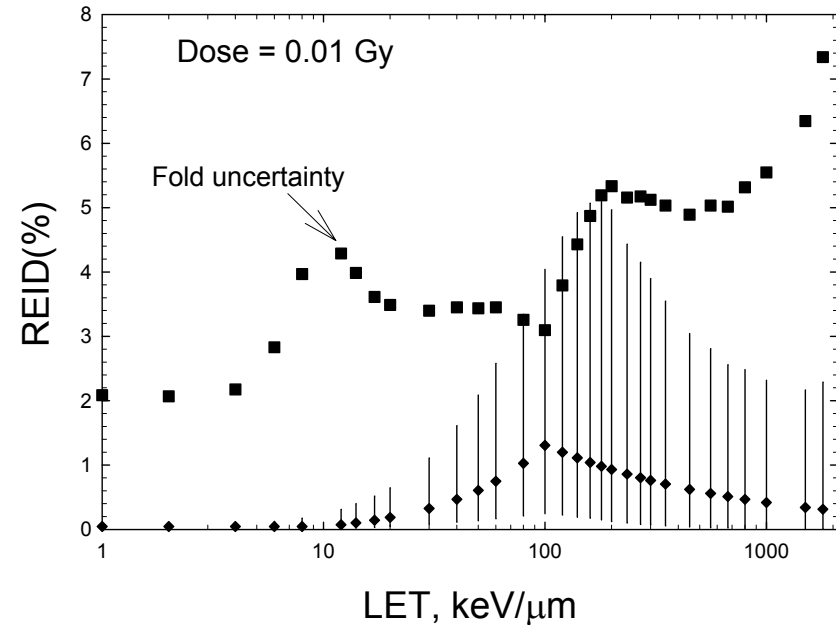


Risk and Uncertainties increase with Linear Energy Transfer (LET)

TEPC Data on NASA-MIR 7 (1/21/1998 to 5/28/1998)



95% CL of Risk Model



$$\text{Dose} = \text{Flux} \times \text{LET}$$

$$\text{Biological Dose } H = \text{Dose} \times Q(L)$$

$$\text{REID} = \text{Risk of exposure induced death} = H \times R_0(\text{sex, age})$$

Accuracy of Space Radiation Transport Models

Absolute Predictions from HZETRN and Flight Measurements

Mission	DATE	Inclination	Altitude	Shielding	Dose, mGy/d			Dose Eq., mSv/d		
					Measured	Theory	%Difference	Measured	Theory	%Difference
STS-40	1991	39	293	Dloc2	0.052	0.048	7.7	0.13	0.16	-23.1
STS-49	1992	28.5	358	Dloc2	0.05	0.048	4.0	0.127	0.155	-22.0
STS-51	1993	28.5	296	Payload Bay	0.044	0.048	-9.1	0.144	0.154	-6.9
STS-57	1993	57	298	Payload Bay	0.113	0.109	3.5	0.422	0.434	-2.8
STS-57	1993	57	298	DLOC-2	0.138	0.11	20.3	0.414	0.37	10.6
Mir-18	1995	51.6	390	P	0.142	0.141	0.7	0.461	0.526	-14.1
STS-81	1997	51.6	400	0-sphere	0.147	0.135	8.2	0.479	0.521	-8.8
STS-81	1997	51.6	400	Poly 3-in	0.138	0.138	0.0	0.441	0.400	9.3
STS-81	1997	51.6	400	Poly 5-in	0.129	0.118	8.5	0.316	0.368	-16.5
STS-81	1997	51.6	400	Poly 8-in	0.128	0.113	11.7	0.371	0.323	12.9
STS-81	1997	51.6	400	Poly 12-in	0.116	0.111	4.3	0.290	0.298	-2.8
STS-89	1998	51.6	393	0-sphere	0.176	0.148	15.8	0.561	0.614	-9.4
STS-89	1998	51.6	393	Al 3-in	0.167	0.159	4.8	0.445	0.488	-9.7
STS-89	1998	51.6	393	Al 7-in	0.149	0.161	-8.1	0.529	0.617	-16.6
STS-89	1998	51.6	393	Al 9-in	0.171	0.162	5.3	0.492	0.541	-10.0

Model vs. Phantom Expt. (STS and ISS)

Phantom Data on STS-91 for Trapped + GCR (51.6 x 390 km)					
Organ	Measured (mGy)	Theory (mGy)	Theory* (mGy)	% Difference	% Difference*
Brain	2.23	2.42	2.26	-8.5	-1.4
Bone Surface	2.16	2.36	2.21	-9.3	-2.1
Esophagus	1.71	1.79	1.67	-4.7	2.2
Lung	1.92	1.81	1.69	5.7	11.9
Stomach	2.05	2.08	1.94	-1.5	5.2
Liver	1.88	2.15	2.01	-14.4	-6.9
Spinal Column	1.65	1.98	1.85	-20.0	-12.1
Bone Marrow	1.75	1.98	1.85	-13.1	-5.7
Colon	1.71	1.9	1.78	-11.1	-3.8
Bladder	1.58	1.87	1.75	-18.4	-10.6
Gonad	1.75	1.85	1.73	-5.7	1.2
Skin/Breast	2.46	2.58	2.41	-4.9	2.0
Skin/Abdomen	2.35	2.58	2.41	-9.8	-2.6

*Includes a correction to TLD efficiency vs. LET.

CALCULATIONS & % DIFFERENCES

	TRAPPED (mGy/day)	GCR (mGy/day)	TOTAL (mGy/day)	% DIFF. (C-M)/M
BRAIN	0.066	0.077	0.143	13.3
THYROID	0.072	0.077	0.148	9.4
HEART	0.061	0.077	0.138	6.6
STOMACH	0.057	0.077	0.133	5.5
COLON	0.056	0.076	0.131	2.5
LIVER	0.053	0.077	0.130	-4.0

ISS
Increment-2
results

HZETRN Comparisons to GCR Secondary Energy Spectra on STS 48

- Because of Earth Magnetic Cutoff predominantly secondary protons and deuterons are measured (Stringent test of HZETRN Code)

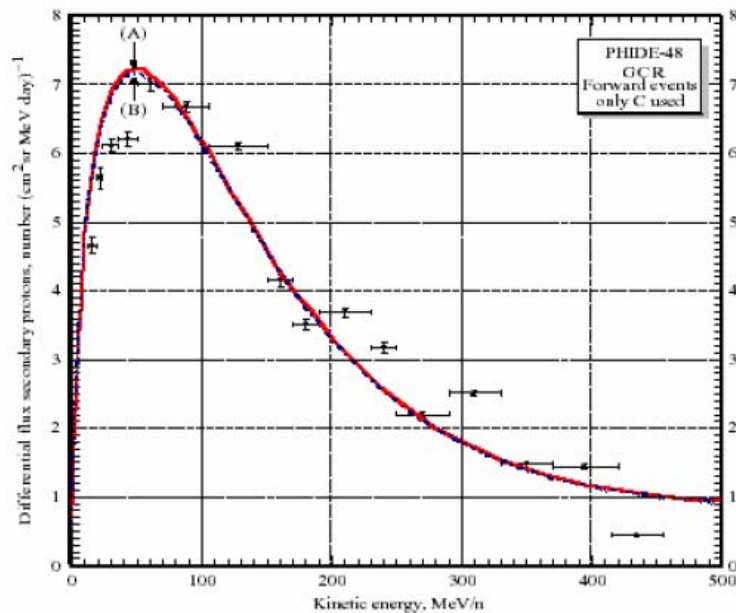


Figure 5. Comparison of observed and calculated secondary proton energy spectra.

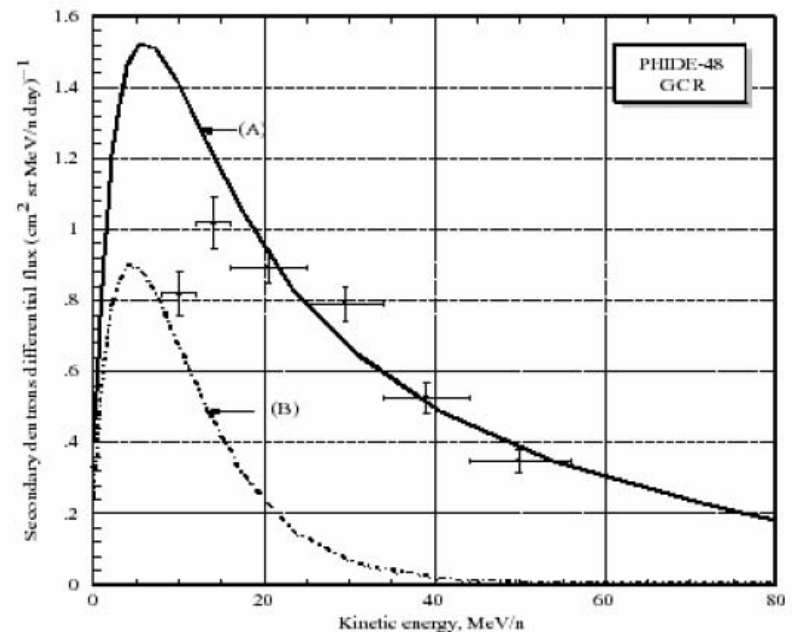
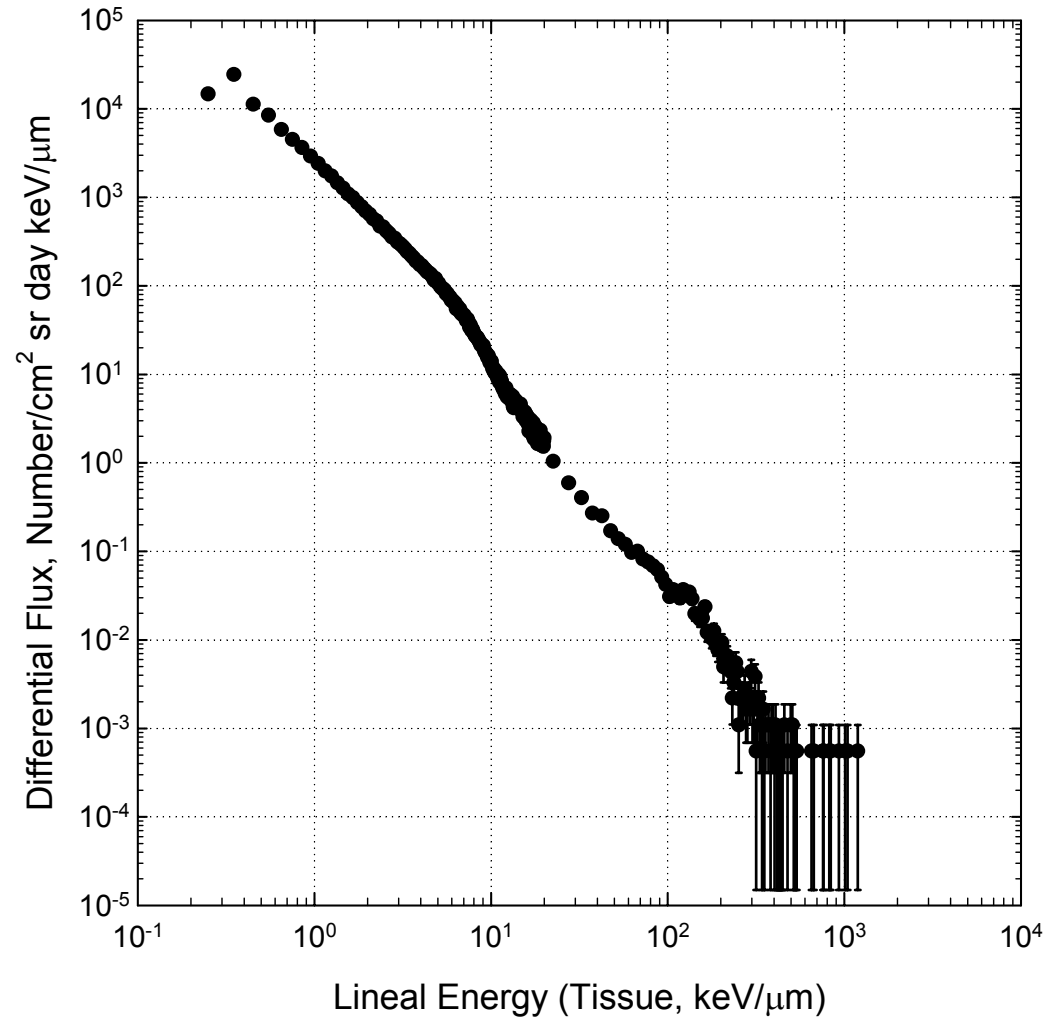


Figure 6. Comparison of observed and calculated secondary deuteron energy spectra.

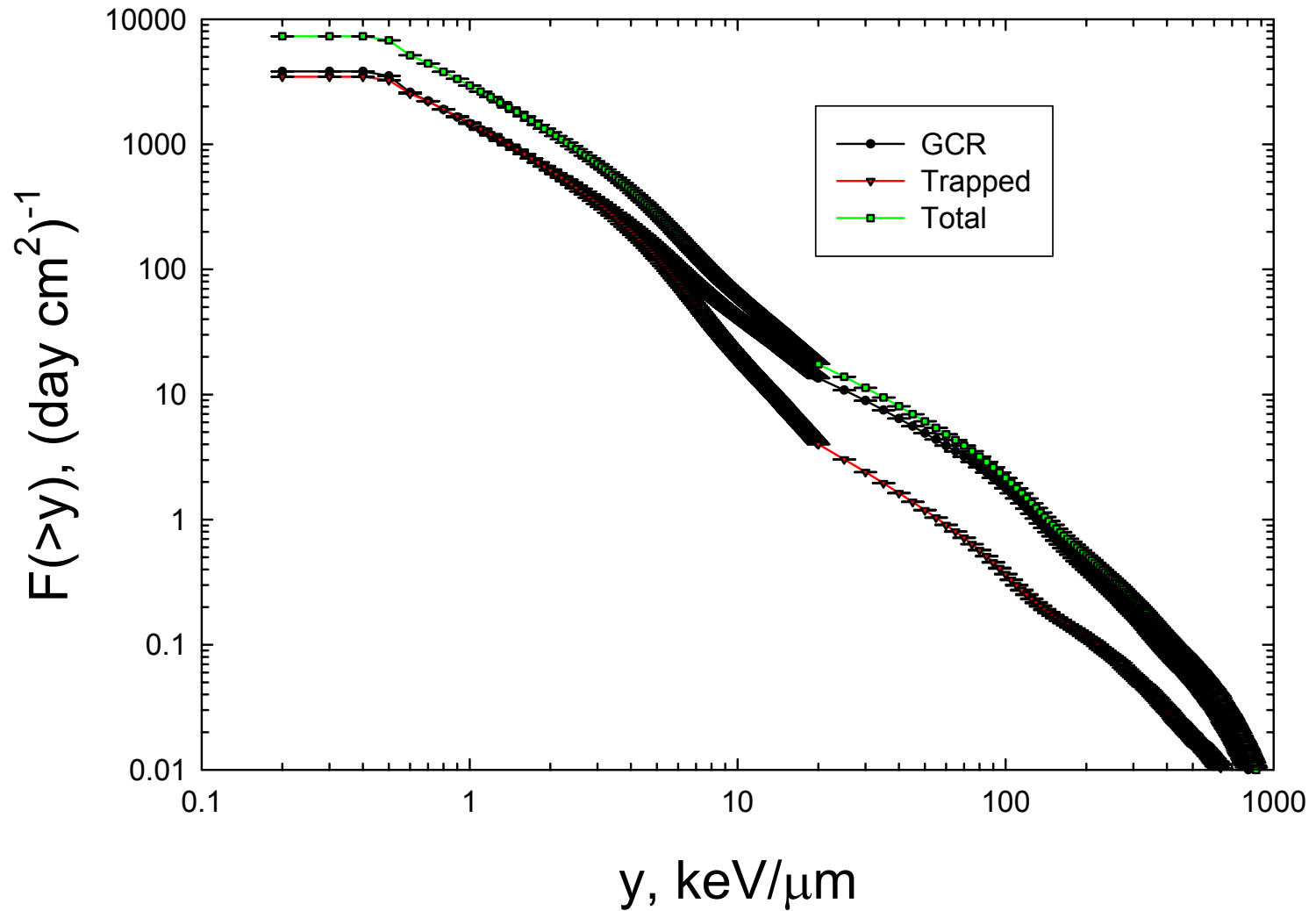


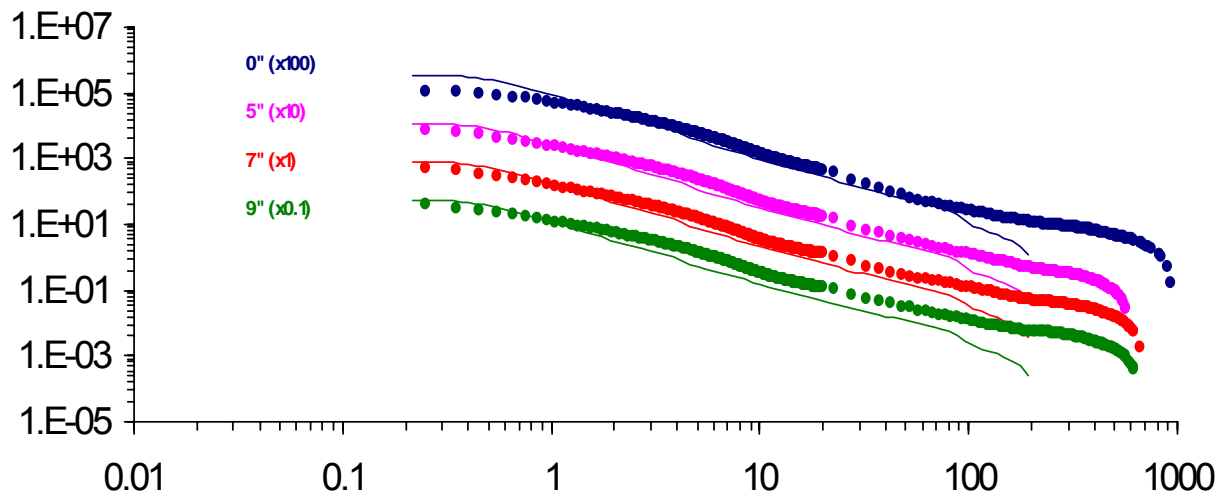
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Phantom Torso TEPC Trapped + GCR Differential Flux June 25 - July 3, 2001

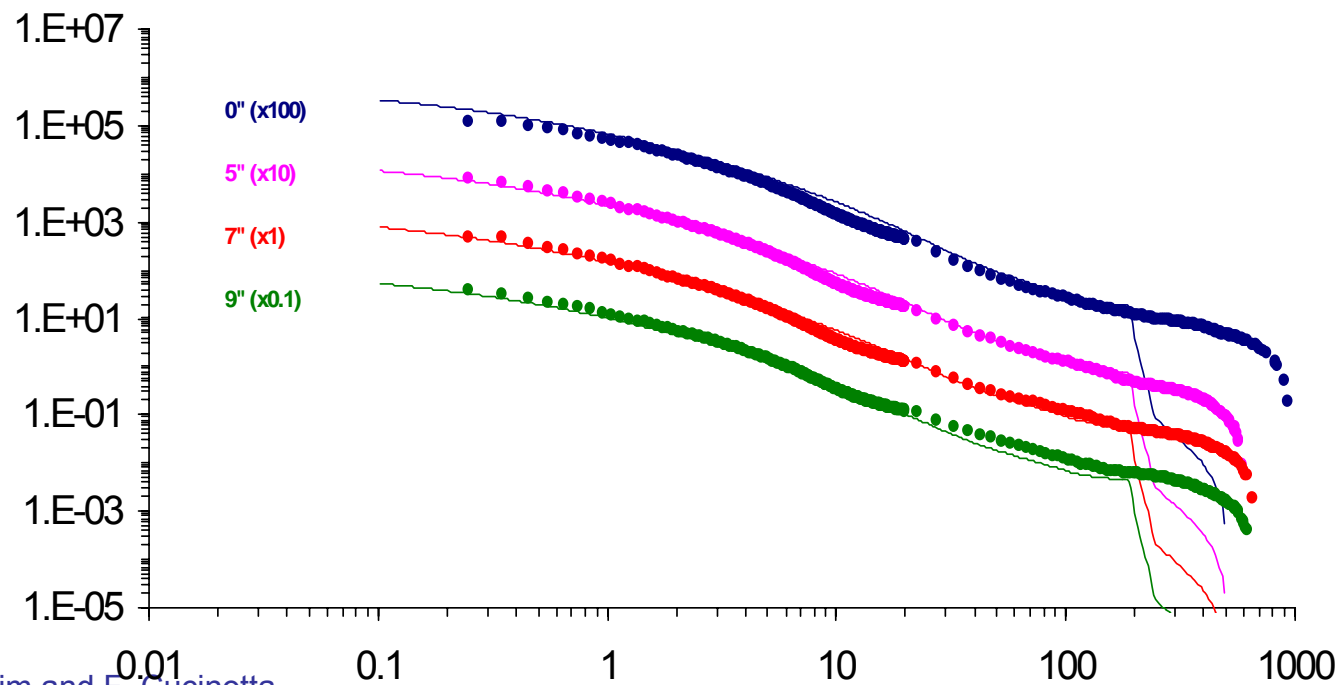


NASA-MIR 7 (1/21/1998 to 5/28/1998)





LET vs TEPC



TEPC model
Vs TEPC (no TF)

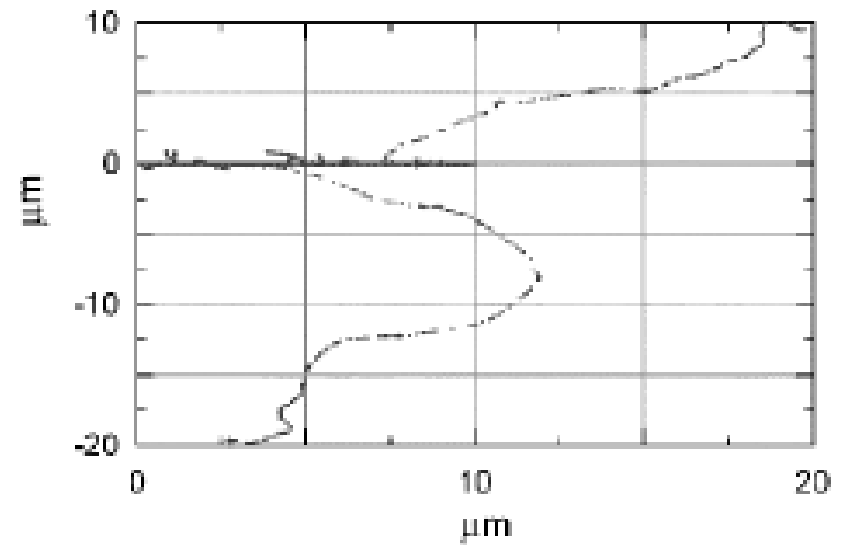
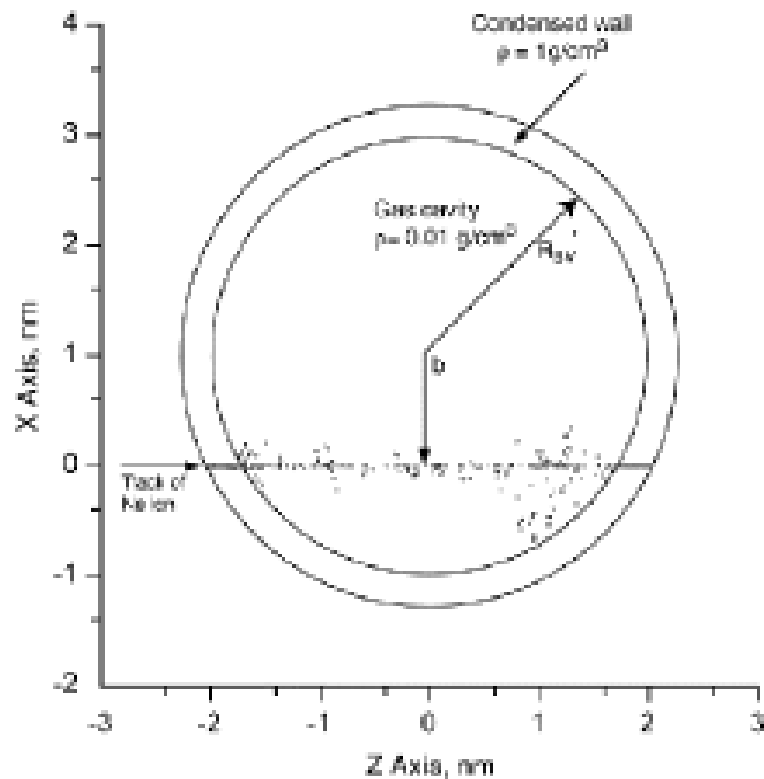


Figure 3. An example of the energy transfer points produced by the passage of a high-energy Ne ion through a walled proportional-counter chamber (left) and of the two-dimensional spatial distribution of two long-range δ -rays (electrons) originating from the primary track (right).